The General Antiparticle Spectrometer (GAPS) -Hunt for dark matter using low energy antideuterons

Identification of Dark Matter - July 2010 Montpellier

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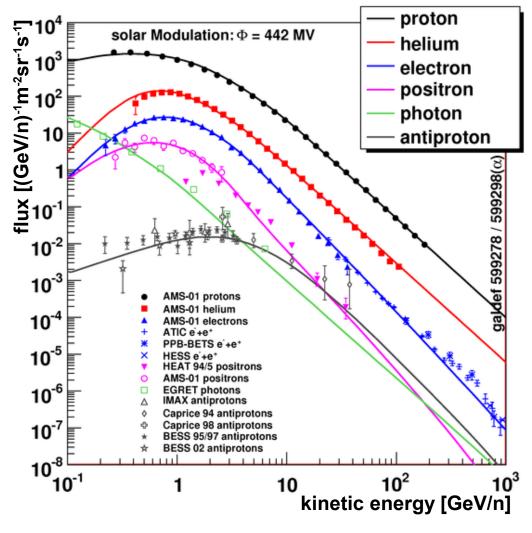
Collaboration

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- 1 UC Berkeley2 Columbia University
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- 5 University of Latvia
- 6 UC Los Angeles



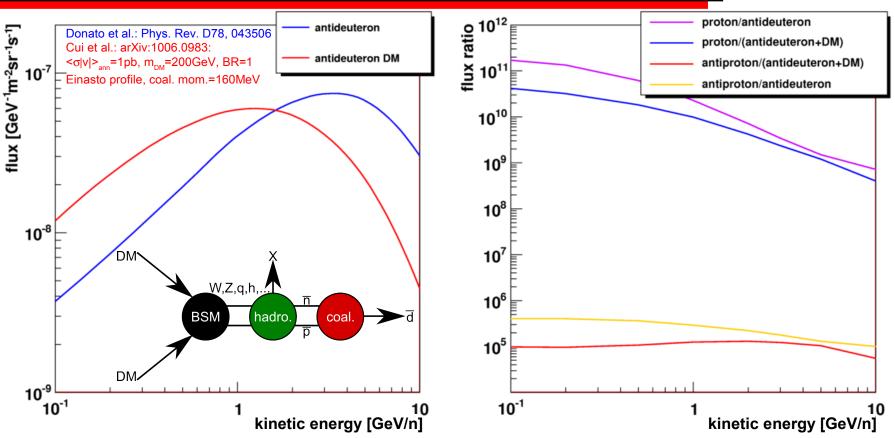
Cosmic rays



what we already learned:

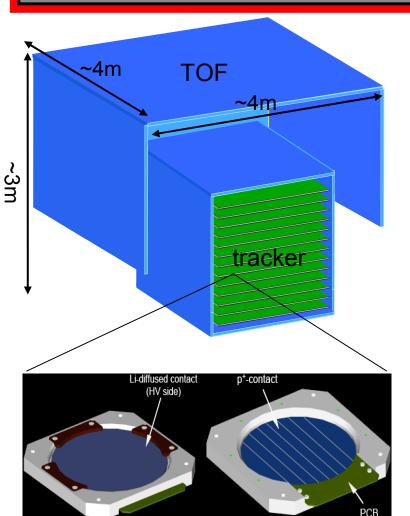
- particle physics
- interstellar medium
- γ-astronomy
- ...
- in general good agreement of models
- search for new phenomena
 - dark matter
 - astrophysical objects
 - baryogenesis
 - unexpected things?

Antideuterons and dark matter



- antideuteron flux is very small.
 - challenging to measure the first antideuterons in cosmic rays (secondary interactions of protons with interstellar gas)
 - good source to study new phenomena
- theories with dark matter self-annihilation predict "large" low-energy antideuteron signals

GAPS concept



Balloon flights planned from 2014

GAPS consists of two detectors (accep.: ~2.7m²sr):

Si(Li) tracker:

- Si(Li) tracker:13 layers composed of Si(Li) wafers
- relatively low Z material (2/3mm,escape fraction ~20keV)
 →target and detector
- Lithium doped Silicon detectors for a good x-ray resolution
- circular modules segmented into 8 strips, ~8cm² each
 →3D particle tracking
- 270 per layer (total: ~3500)
- timing: ~50ns
- dual channel electronics
 5-200keV: X-rays (resolution:~2 keV)
 0.1-200MeV: charged particle

Time of flight and anticoincidence shield:

- plastic scintillator with PMTs surrounds tracker
- track charged particles
- velocity measurement
- anticoincidence for charged particles

Designed for low-energy antideuterons:

no need for a magnet (heavy, sometimes complicated/expensive)

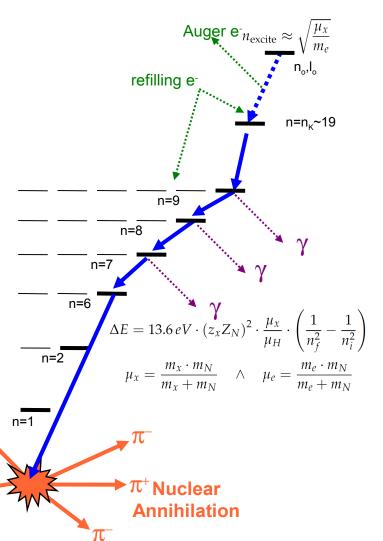
Antideuteron identification

- antideuteron slows down and stops in material
- large chance for creation of an excited exotic atom (E_{kin}~E_I)
- deexcitation:
 - fast ionisation of bound electrons (Auger)
 - complete depletion of bound electrons
 - Hydrogen-like exotic atom (nucleus+antideuteron) rad. deexcitation:

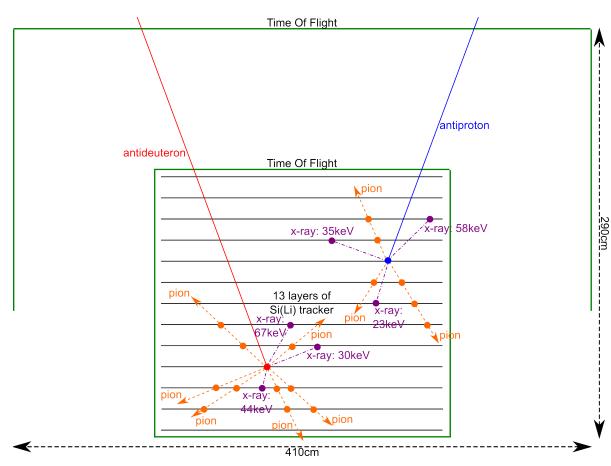
characteristic x-ray transitions

- nucleus-antideuteron annihilation: pions
- exotic atomic physics quite well understood (tested in KEK 2004 testbeam)

atomic transitions



Backgrounds



GAPS needs a very reliable particle identification:

- •identification uses:
 - TOF velocity and tracks
 - depth in tracker
 - x-rays and pions from annihilation
- •important background sources
 for antideuteron events:
 - antiprotons
 - protons, electrons, neutrons in coincidence with cosmic x-rays
 - atmospheric production of antideuterons
 - etc...needs to studied in great detail

GAPS antideuteron sensitivity

- GAPS need small geomagnetic cut-off
 - → therefore (ultra) long duration balloon flights from South Pole are planned: 60 (300) days
- different scenarios give reasonable antideuteron fluxes within sensitivity:

Supersymmetry:

LSP: neutralino (majorana)

Kaluza-Klein UED:

LKP: 1st excitation of photon (boson)

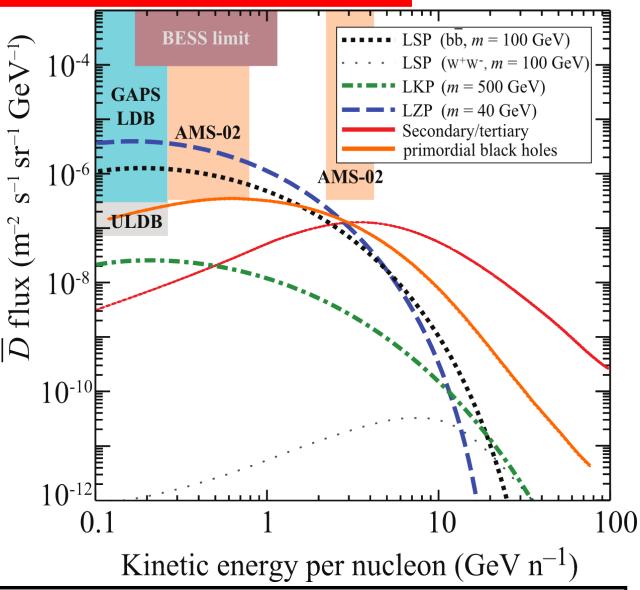
Warped extra dimensions:

Dirac fermion with flavor dependent couplings to Z' primordial black holes:

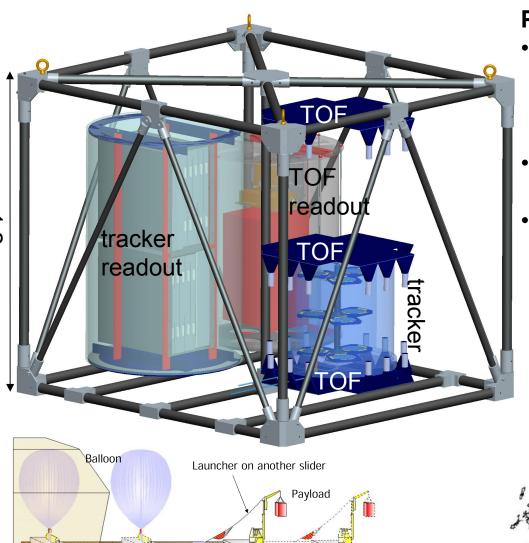
flux from Hawking radiation

 synergy with direct searches and neutrino telescopes:
 GAPS covers complementa

GAPS covers complementary dark matter regions!



Prototype experiment



Prototype GAPS (pGAPS) goals:

- demonstrate stable, low noise operation of the detector components at float altitude and ambient pressure.
- demonstrate the Si(Li) cooling approach and verify thermal model
- measure incoherent background level in a flight like configuration.

Tokyo 2011 scheduled flight from Taiki, Japan

Sanriku

Spooler on a slider

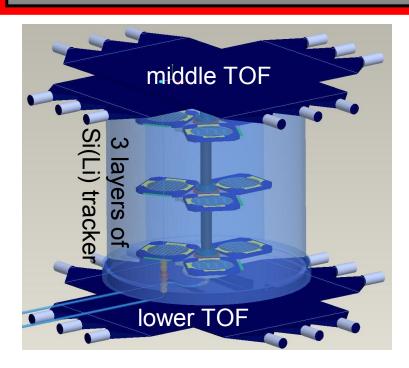
Taiki Aerospace Research Field

E 143° 26′

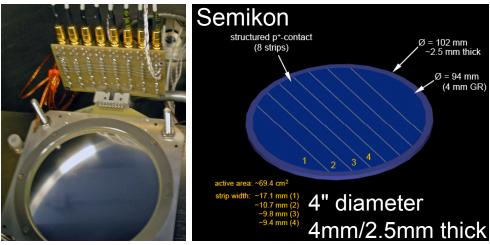
Balloon Center

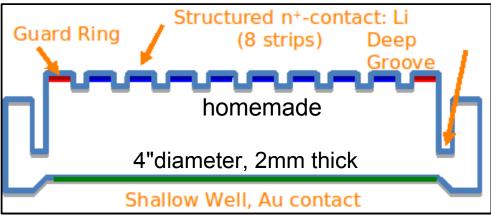
30' 00"

Si(Li) tracker



- 7 commercial Semikon detectors
- 2 homemade detectors (test for the bGAPS fabrication)
- energy resolution at 60keV should be 3keV
- operation at ambient pressure (8mbar)
- cooling system delivers: -35°C





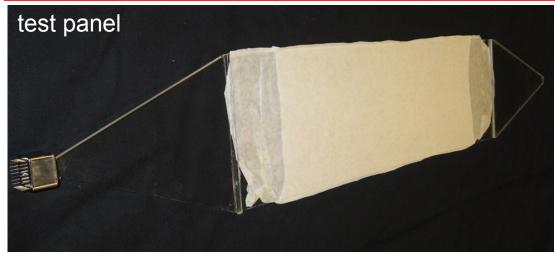
Semikon: N+: Lithium contact

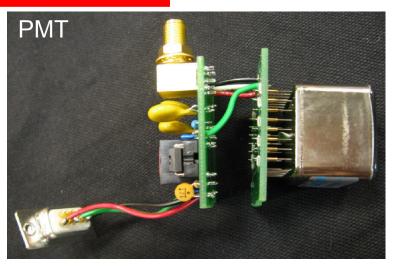
P+: Boron implanted (strips)

Homemade: N+: Lithium contact (strips)

Au contact with shallow well

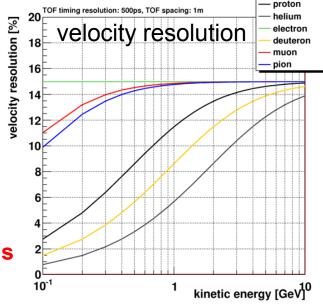
Time of flight system



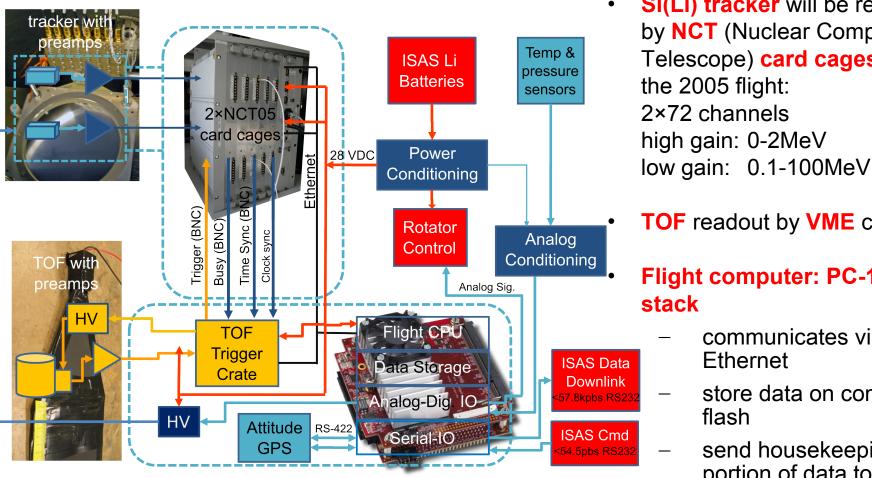




- 3 planes of TOF1 plane consists of 3×3 crossed panels1 panel has 2 PMTs
 - = 18 panels and 36 PMTs
- or Bicron (BC-408)
- Hamamatsu R-7600 PMT
- timing resolution: 500ps
- charge resolution: 0.35e
- MOP value: ~15 photo electrons
- angular resolution: 8°



Readout



Si(Li) tracker will be read out by NCT (Nuclear Compton Telescope) card cages from the 2005 flight: 2×72 channels high gain: 0-2MeV

TOF readout by **VME** crate

Flight computer: PC-104 stack

- communicates via **Ethernet**
- store data on compact flash
- send housekeeping, portion of data to ground
- commanding

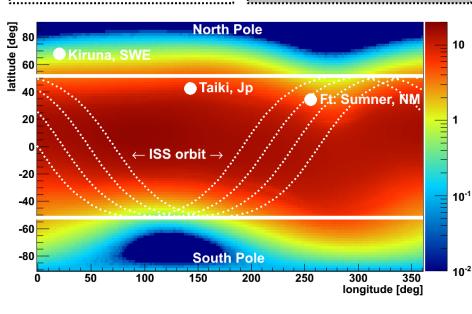
Simulation

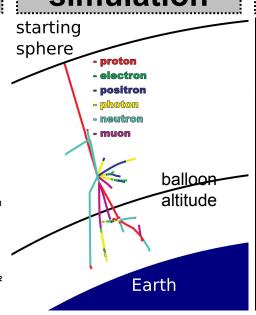
cosmic antideuterons

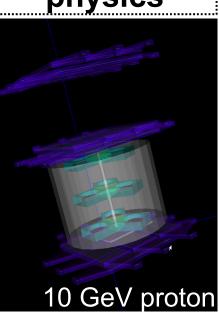
atmospheric simulation

detector simulation

exotic atomic physics







Atmospheric and geomagnetic simulations with PLANETOCOSMICS based on GEANT 4.9.2

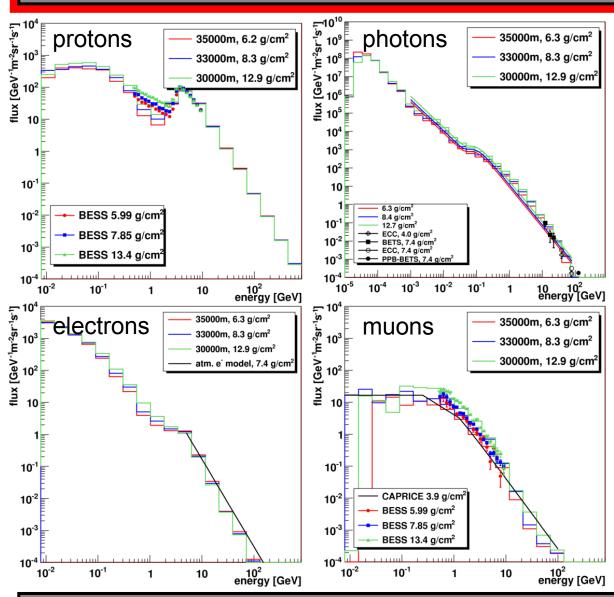
- geomagnetic simulations as a function of position and direction
- atmospheric fluxes is in good agreement with measurements

Instrument simulation with GEANT, ROOT output format:

- electromagnetic, hadronic, optical physics are running
- basic components are implemented, frames and structures must be added
- ion and exotic physics are under development

more details at Geant4 Space Users Workshop 2010. Seattle

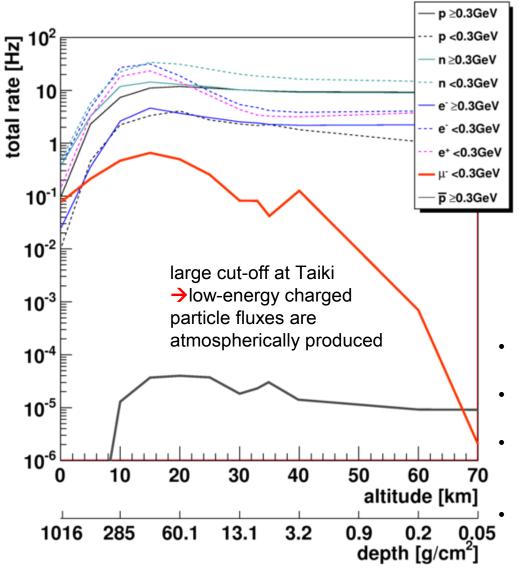
Validation of simulation

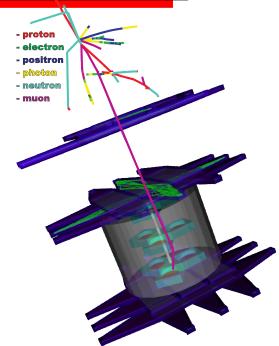


- particle fluxes

 (ATM+CR) for certain
 particle types at
 different altitudes
- comparison of atmospheric simulations shows good agreement with BESS, ECC, BETS, PPB-BETS, CAPRICE measurements and models
- simulations may need check for light ion physics

Particle rates for pGAPS



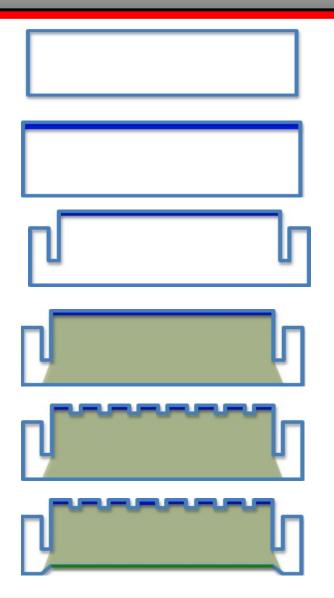


- particle rate at 33km at Taiki: total ~10-50Hz (accep.: 0.054m²sr)
- strongest backgrounds by neutrons, protons, electrons, positrons
- no antiprotons will be measured, but muons might be used to create muonic atoms to study exotic physics at flight
- exotic atomic events faked by x-ray coincidences with cosmic rays in the detector need to be studied **carefully!**

Conclusion & Outlook

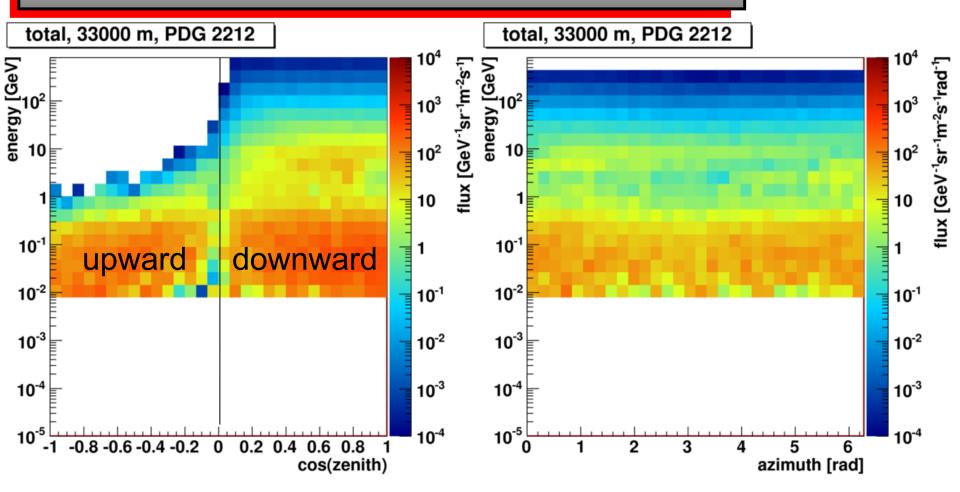
- measurement of low-energetic antideuteron flux is a promising way for indirect dark matter search (probably even more than positrons and antiprotons)
- GAPS is specifically designed for low-energetic antideuterons with a unique detection technique using the creation of exotic atoms
- GAPS is planned to have (U)LDB flights from South Pole starting from 2014
- prototype experiment is currently under construction and a flight is scheduled for Summer 2011 from Taiki, Japan:
 - hardware development: Si(Li), TOF, readout, structure, thermal model
 - simulation of particle fluxes in atmosphere and detector

Si(Li) fabrication



- 1. Cut from the ingot
- 2. Evaporate Lithium
- 3. Produce the deep groove and mesa (optional)
- 4. Drift the Li into the silicon
- 5. Make strips and guard ring
- 6. Etch the back (shallow well) and evaporate Au

Proton fluxes at Taiki



- total proton fluxes (cosmic + atmospheric) at 33km altitude
- upward fluxes have smaller energies
- no dependence on azimuth angle