

xBSM data from May 2010 run

J.W. Flanagan

CesrTA WebEx

2010.7.13

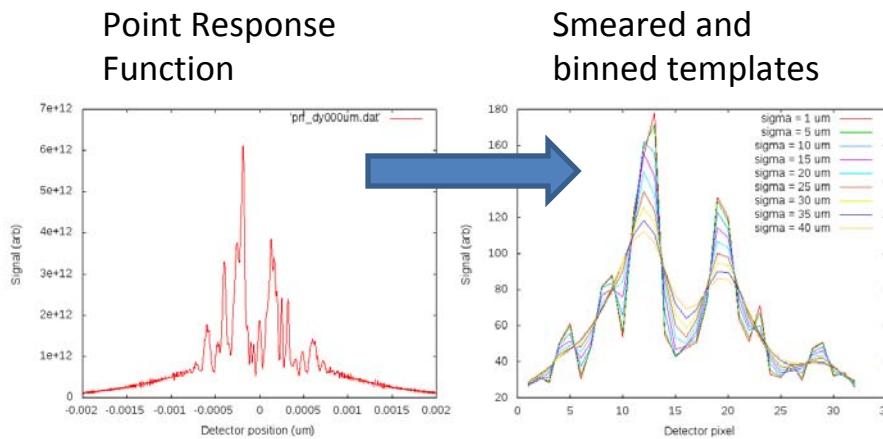
Contents

- E-cloud study data (CA, FZP)
- Resolution estimates and comparison with e-data at D Line 2 GeV
- Resolution estimates for other lines and energies (C and D, 2 and 4 GeV)
- Not: Calibration not yet done, so all data shown are with uncalibrated interpixel gains.

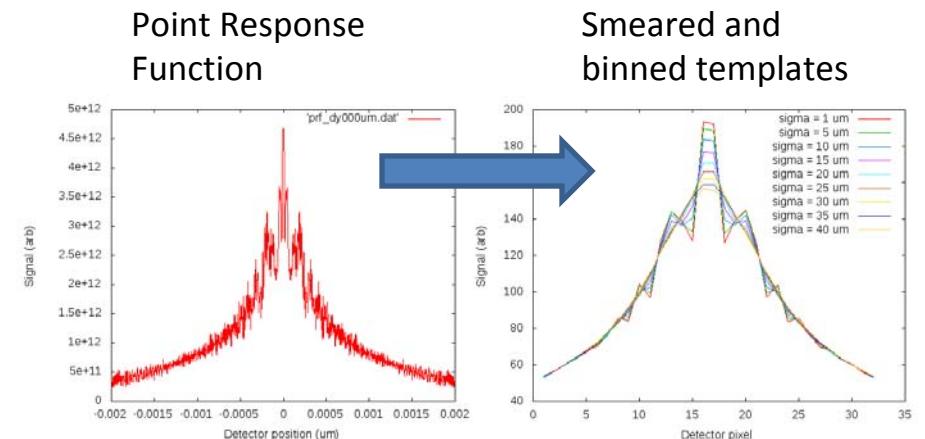
E-Cloud study data

- Data taken night of 10/11 May at D Line (e^+)
- Fill pattern: 45 bunches, 14 ns spacing
- Bunch currents: 0.5, 1.0, 1.15, 1.3 mA
- X-ray optics: CA and FZP
 - FZP data analyzed same way as CA data
 - Fit data to templates generated by sets of point response functions, which are weighted over the assumed source distribution. Grid of templates in y and σ_y used.
 - FZP is treated as being *one-dimensional*. (Cylindrical lens.)

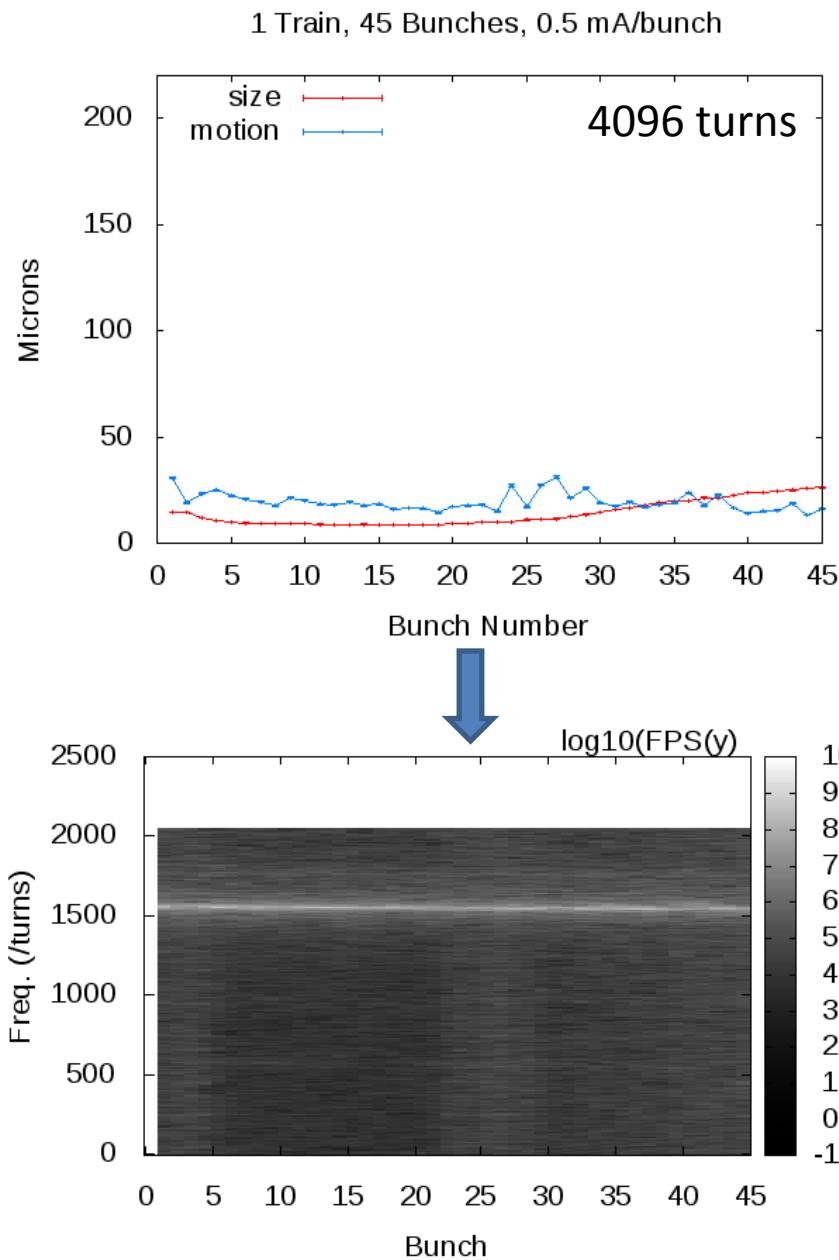
CA



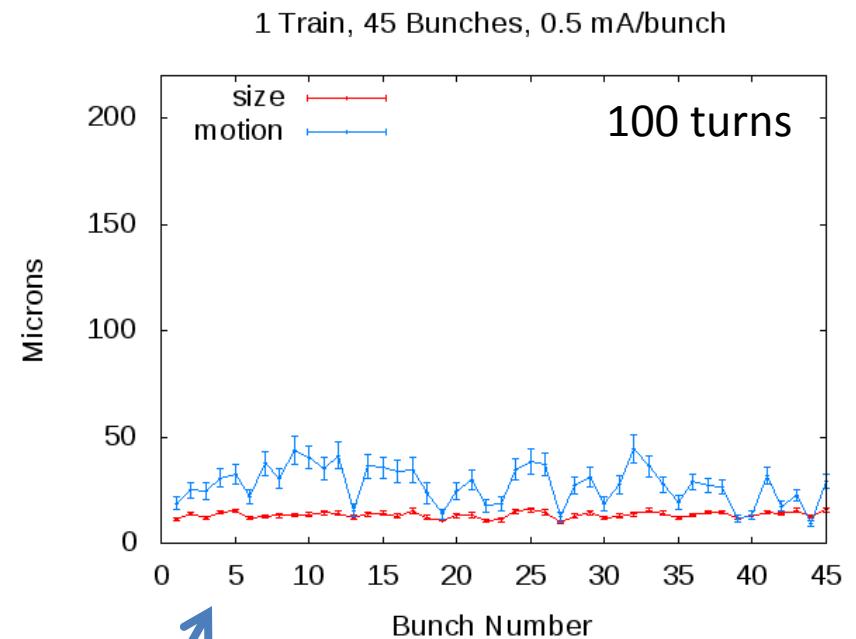
FZP



CA



FZP

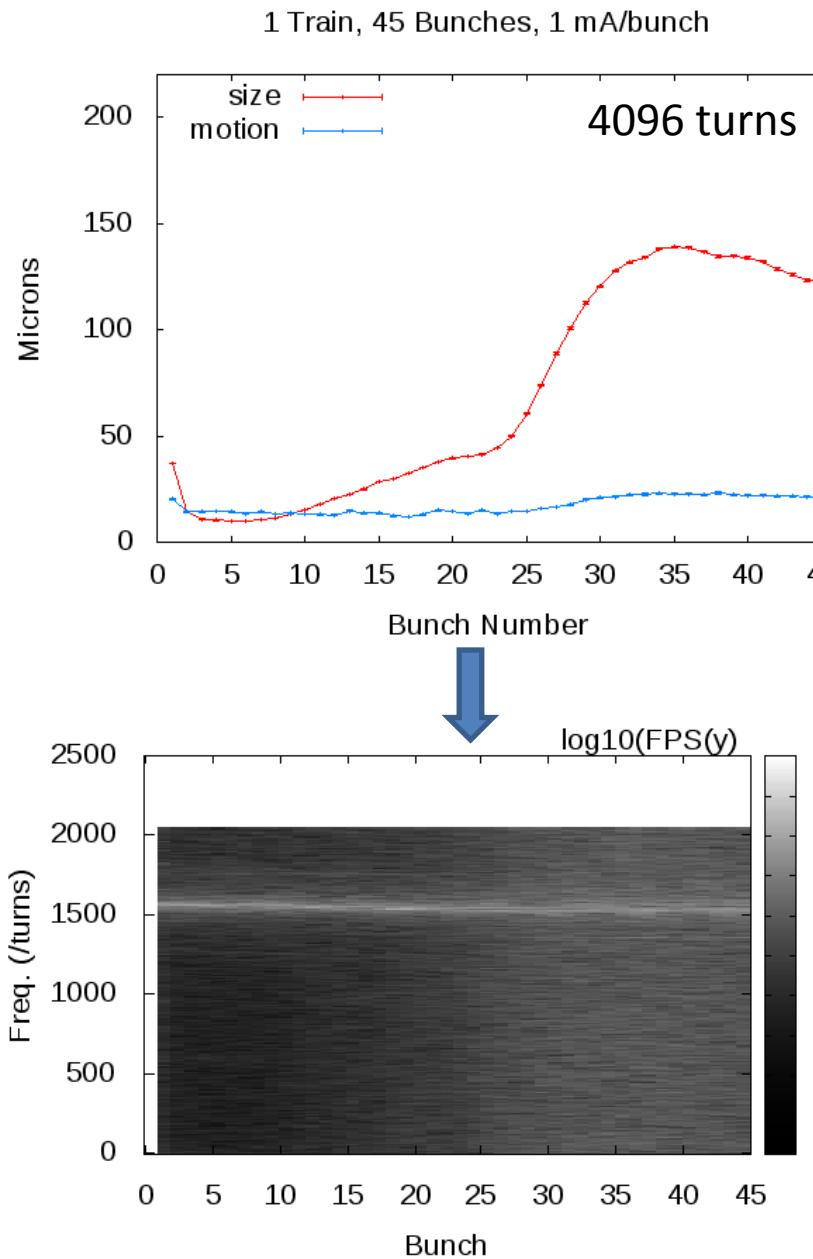


Single-bunch,
single-turn fits
averaged over
turns

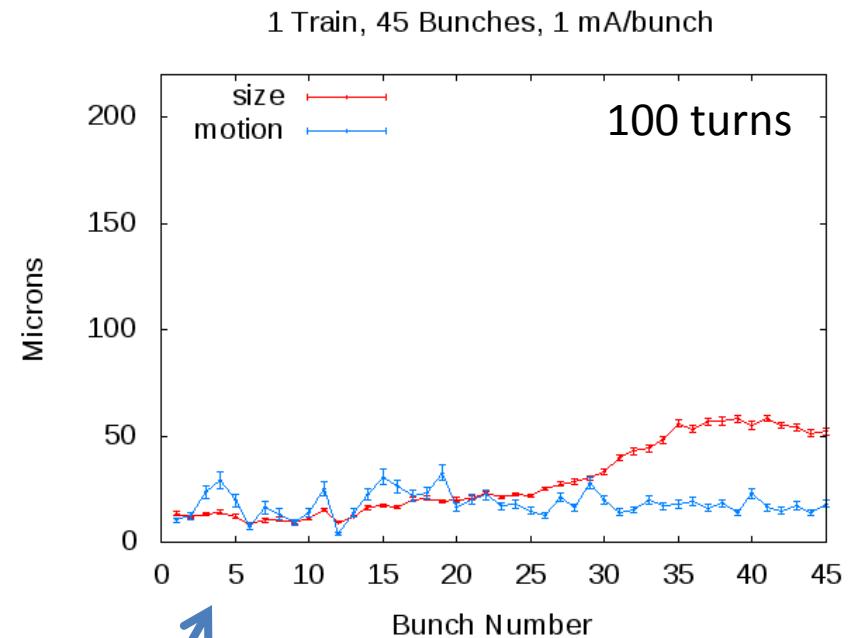
Bunch-by-
bunch
position
spectra

0.5 mA
/bunch

CA



FZP

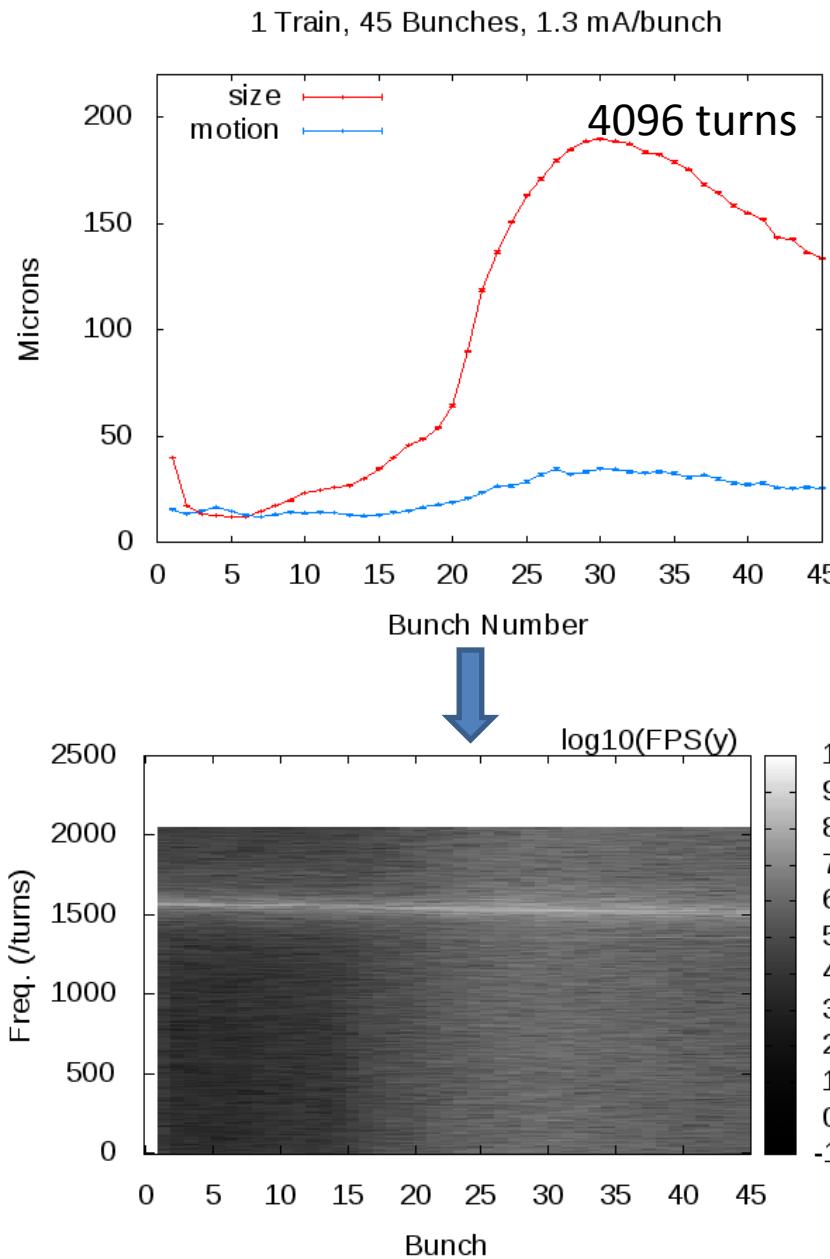


Single-bunch,
single-turn fits
averaged over
turns

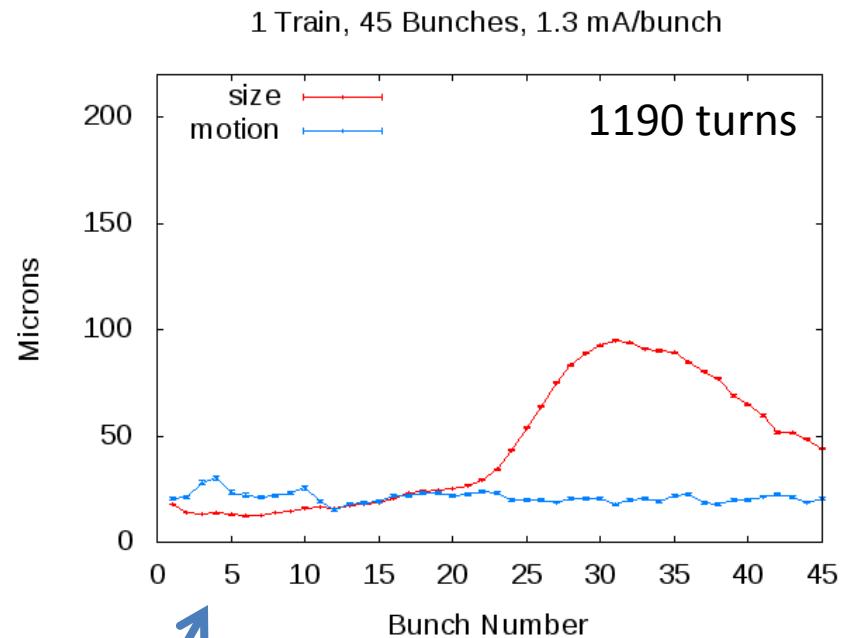
Bunch-by-
bunch
position
spectra

1.0 mA
/bunch

CA



FZP



Single-bunch,
single-turn fits
averaged over
turns

Bunch-by-
bunch
position
spectra

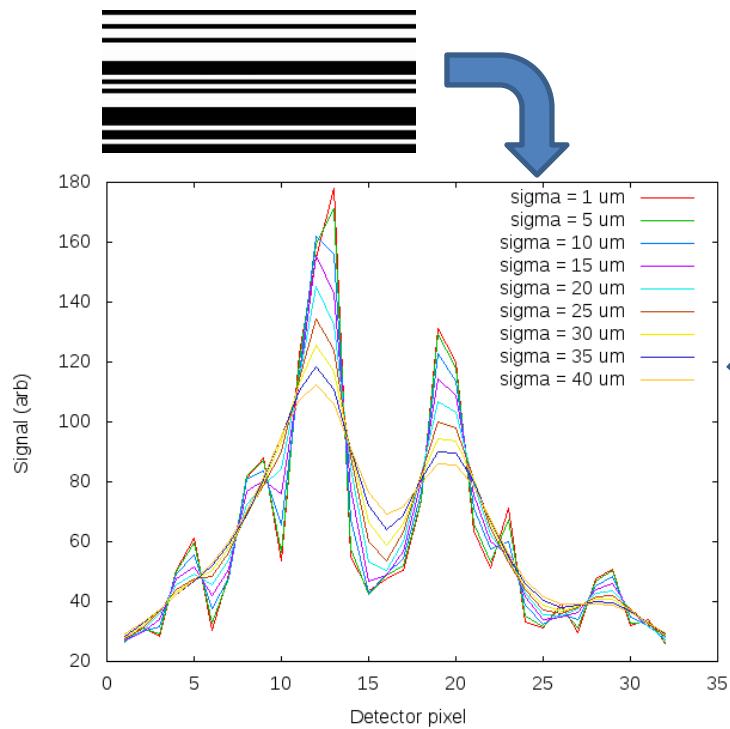
1.3 mA
/bunch

Resolution Estimates

- **Procedure:**
 - Simulate detector images for beams of $\sigma_y = 1\text{-}40 \mu\text{m}$.
 - Fit images against each other. Chi-squared is calculated for cross-fits:
 - $\text{Chisq}/\text{nu} = (1/(N-n-1)) * \text{SUM}[(y_i - y(x_i))^2 / \sigma_i^2]$
 - E.g., 5- μm pattern is checked for fit against 1-, 2-, 3-, ... μm patterns.
 - Bin (pixel) weights are assumed to be statistical ($\sigma_i = \sqrt{y(x_i)}$), assuming average bin height represents 200 photons.
 - Expect ~ 360 photons/pixel/turn/mA/bunch at 2 GeV with low-energy chip
 - Expect ~ 240 photons/pixel/turn/mA/bunch at 4 GeV with high-energy chip
 - Chi-sq 70% exclusion values are taken to represent the resolution contours.
 - Should approximate something like 1-sigma contours.
- **Note that these are *single-shot resolutions*.**
 - **Detector noise is not included, only photon statistics**

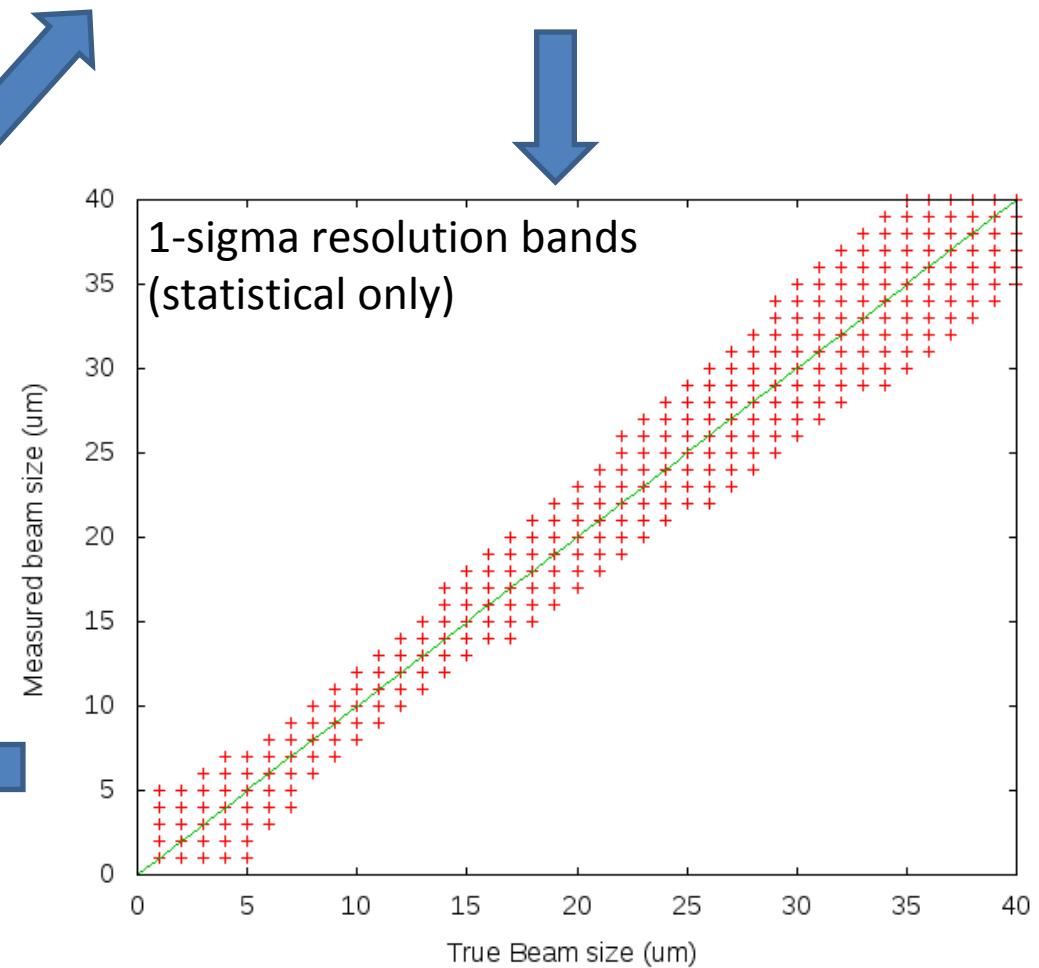
$10 \mu\text{m}$, 31-element CA mask @ D Line 2 GeV

Generate detector images for various beam sizes:



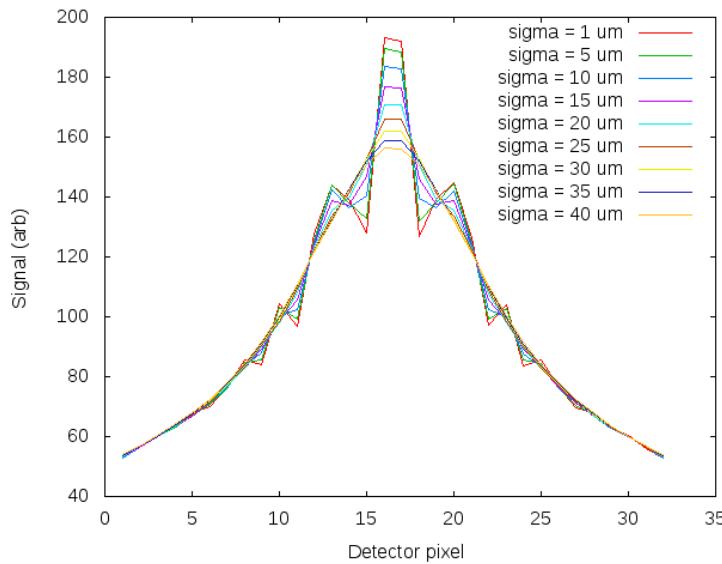
Statistical single-shot resolution at $10 \mu\text{m}$ beam size = $\pm \sim 2 \mu\text{m}$
(Assuming ideal detector.)

Cross-fit between beam sizes.
Plot 1-sigma statistical confidence regions,
Assuming 200 photons/pixel average
($\Rightarrow 0.56 \text{ mA}$ at 2 GeV):



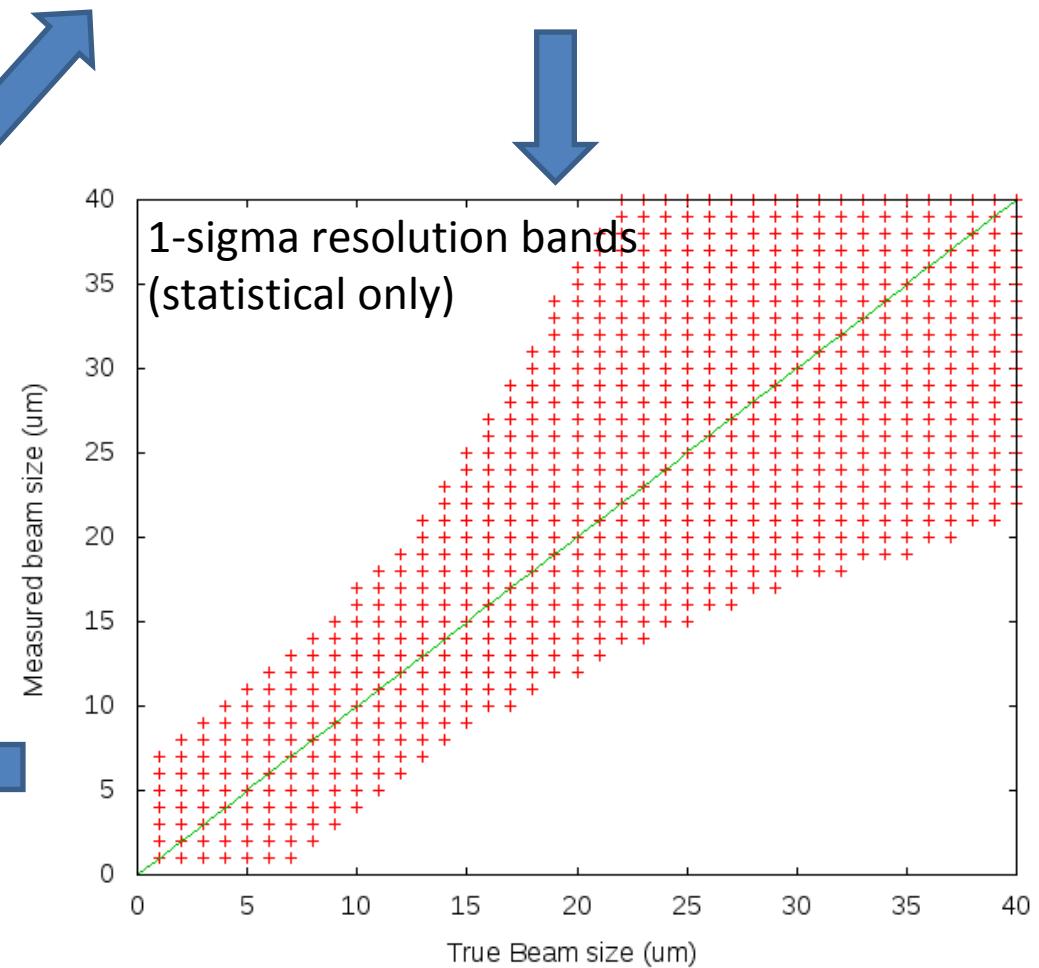
FZP@ D Line 2 GeV

Generate detector images for various beam sizes:



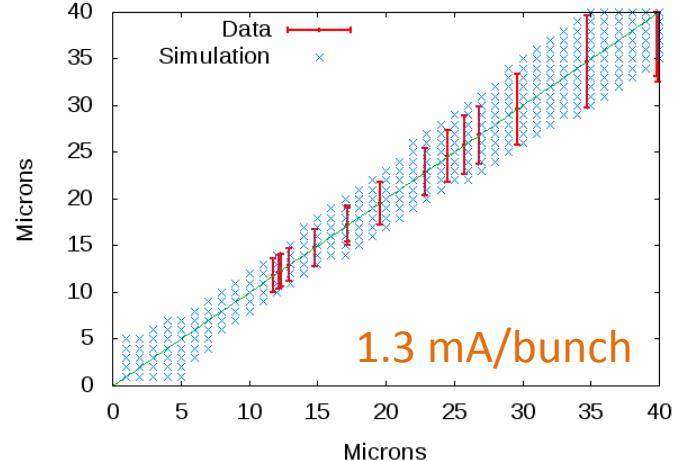
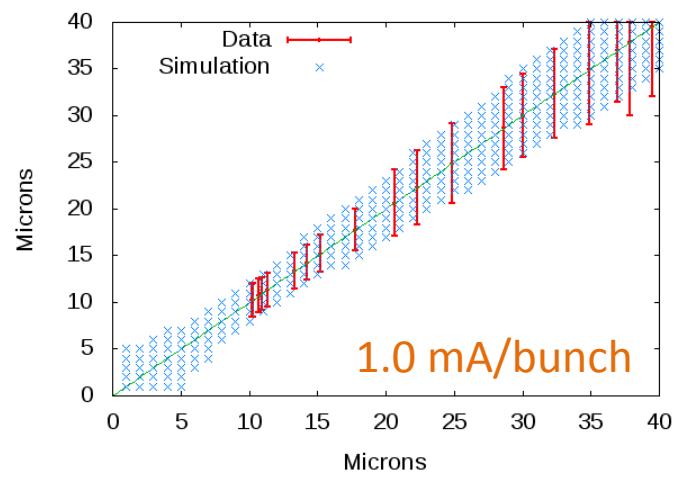
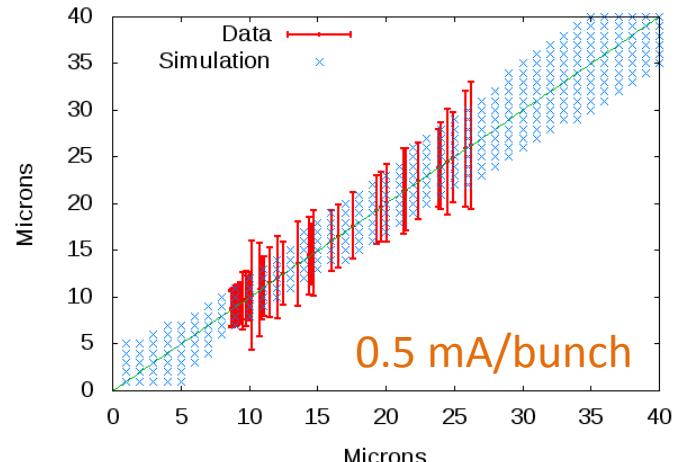
Statistical single-shot resolution at 10 μm beam size = +/- ~6.5 μm
(Assuming ideal detector.)

Cross-fit between beam sizes.
Plot 1-sigma statistical confidence regions,
Assuming 200 photons/pixel average
(=> 0.56 mA at 2 GeV):



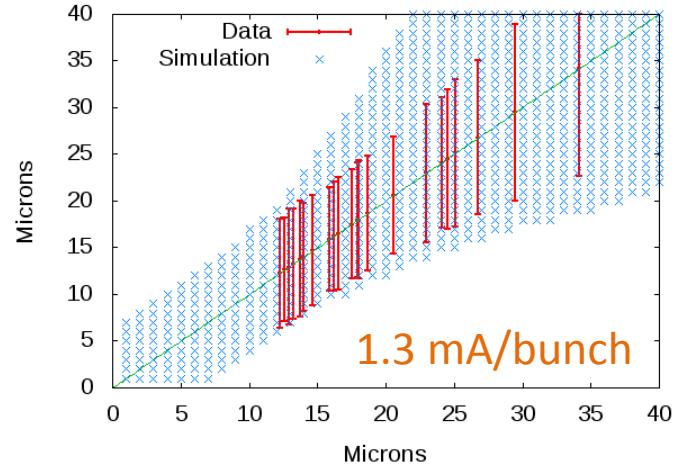
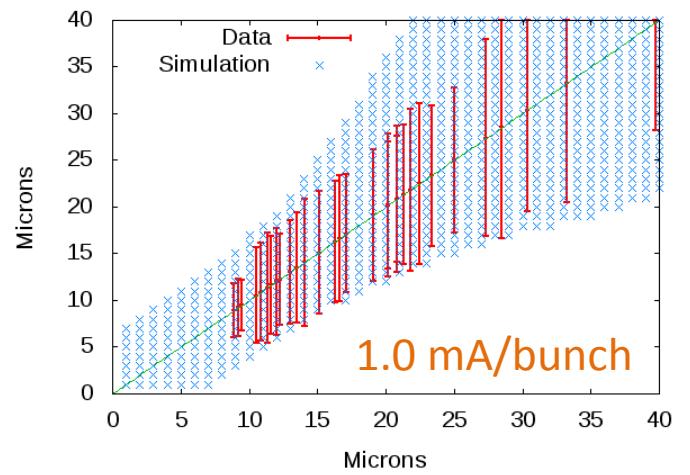
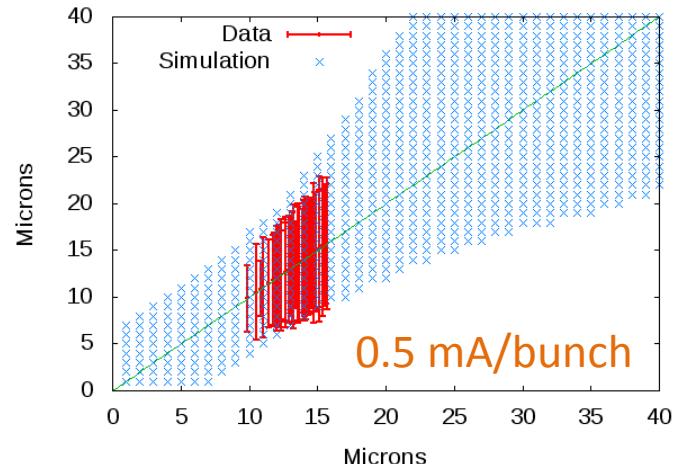
Resolution data vs simulation: CA

- Simulation statistical confidence bands assume
 - Perfect, noiseless detector
 - 200 photons/pixel/shot on average
 - $\Rightarrow 0.56 \text{ mA/bunch}$
- Shot-by-shot spread in data is closest to simulation at nearer to 1 mA.
 - Not using a perfect, noiseless detector.
- Seems reasonable agreement



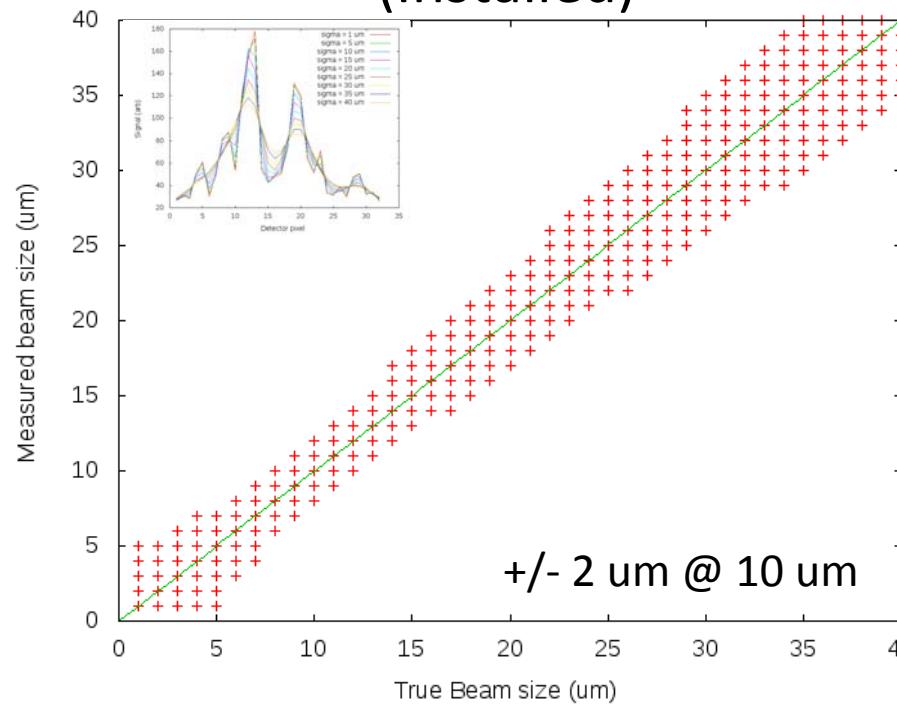
Resolution data vs simulation: FZP

- Simulation statistical confidence bands assume
 - Perfect, noiseless detector
 - 200 photons/pixel/shot on average
 - $\Rightarrow 0.56 \text{ mA/bunch}$
- Not so good agreement at ends:
 - $\geq \sim 18 \text{ um}$: Actual spread smaller than simulation. (?)
 - $< \sim 10 \text{ um}$: Smaller spread due to cutoff in fits?
 - Fits at edge of grid (5 um) are thrown away, and FZP has a lot of “underflow fits” at small beam sizes.

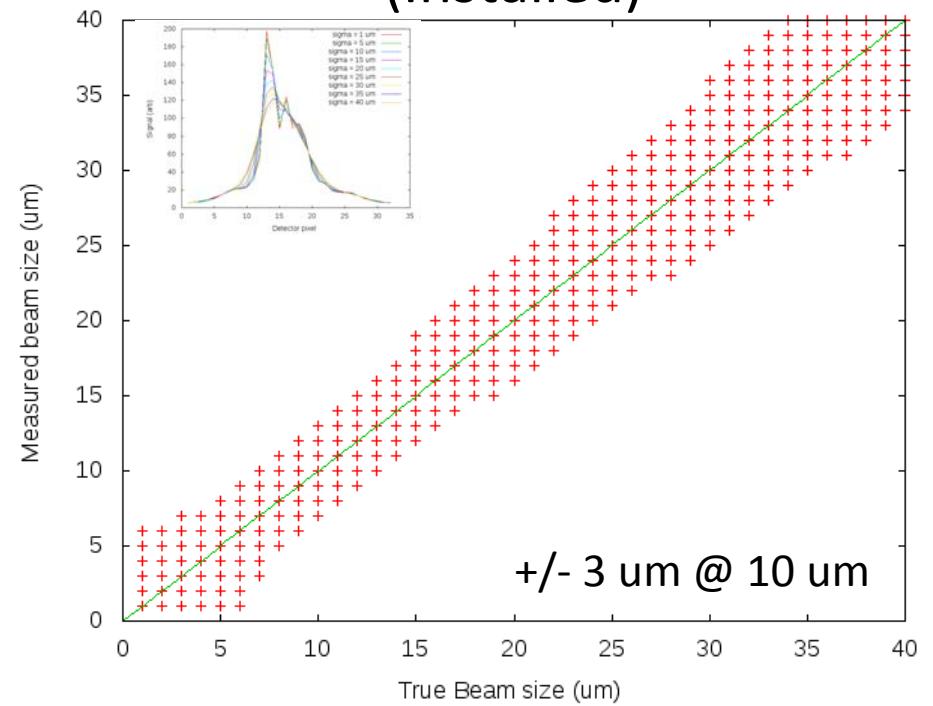


2 GeV CA Resolutions

D Line, 10 μm 31-pixel CA
(Installed)



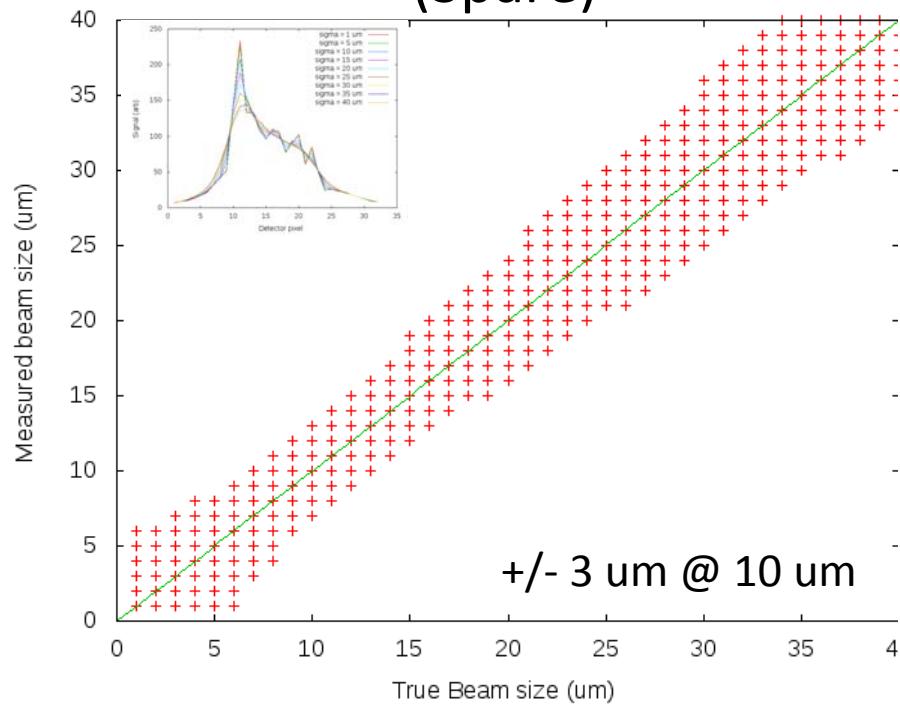
C Line, 5 μm 31-pixel CA
(Installed)



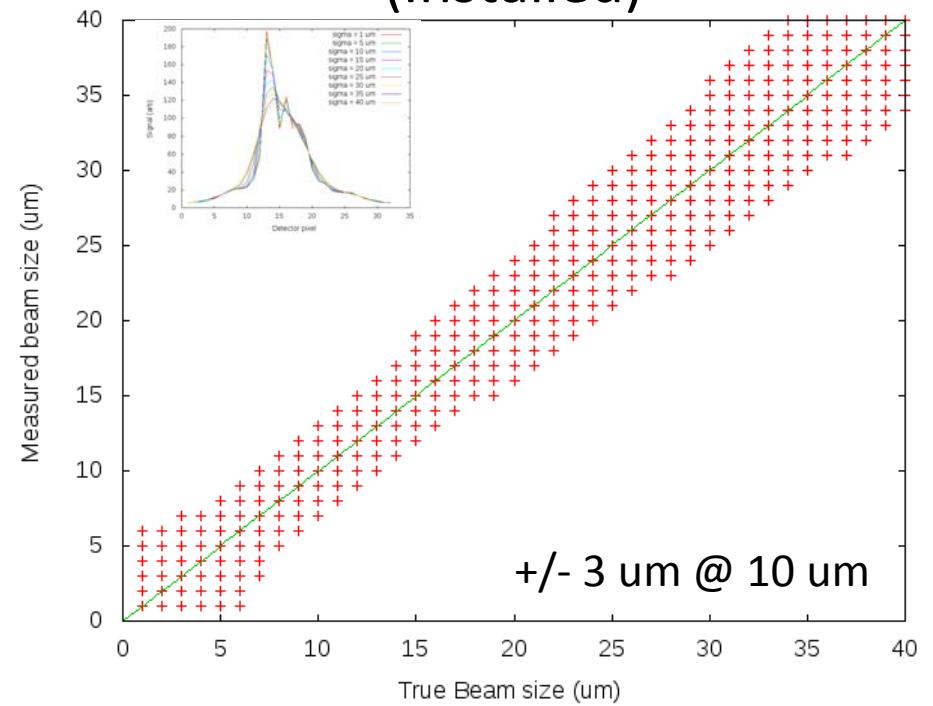
10 μm mask is better...

2 GeV CA Resolutions

C Line, 5 um 47-pixel CA
(Spare)



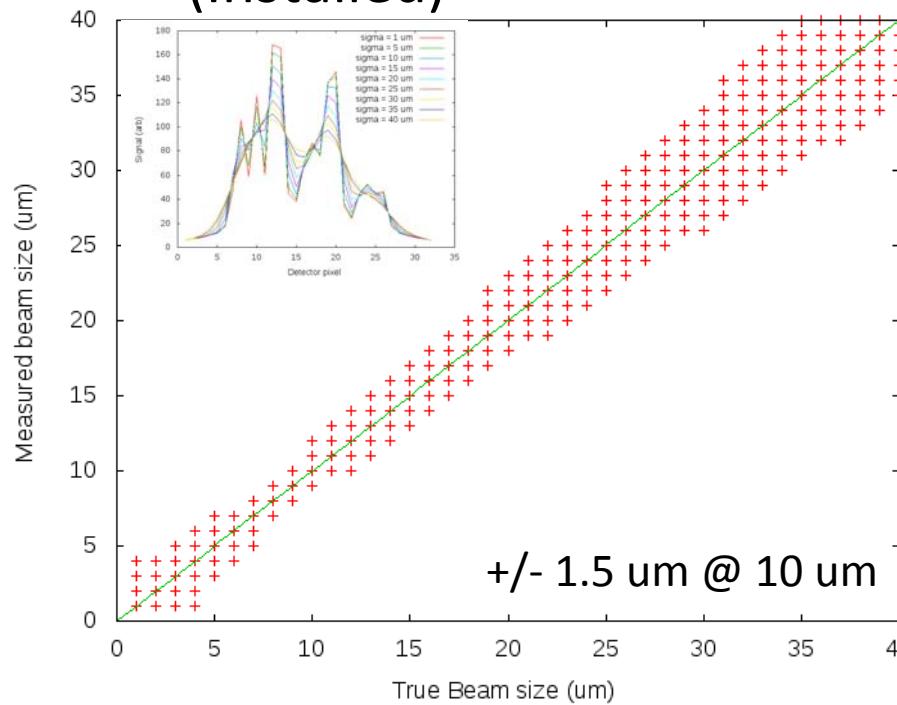
C Line, 5 um 31-pixel CA
(Installed)



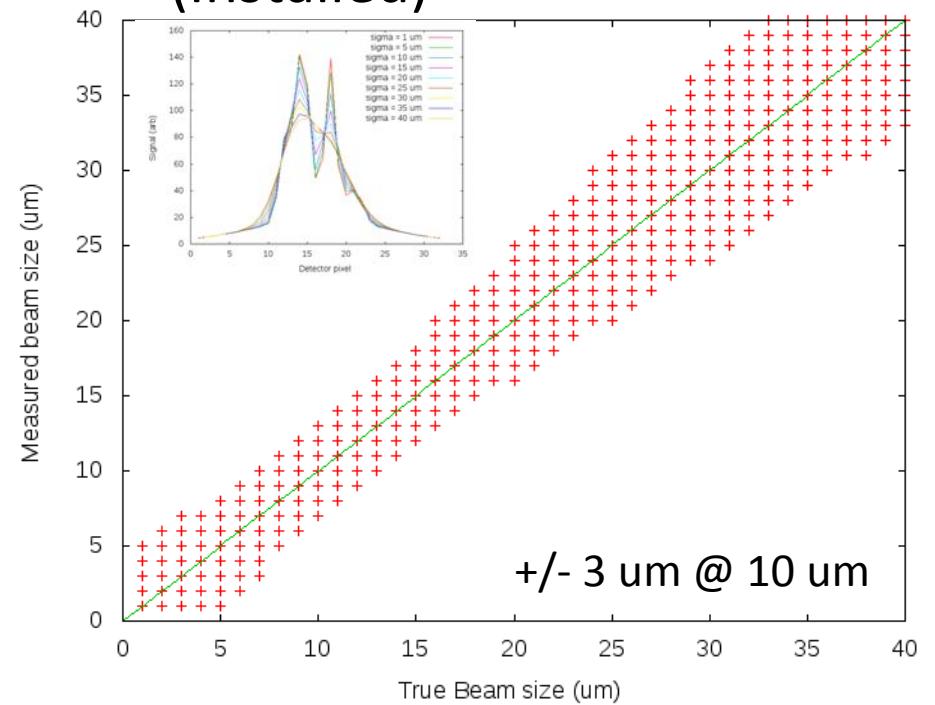
47-pixel mask is no improvement over 31.

4 GeV CA Resolutions

D Line, 10 μm 31-pixel CA
(Installed)



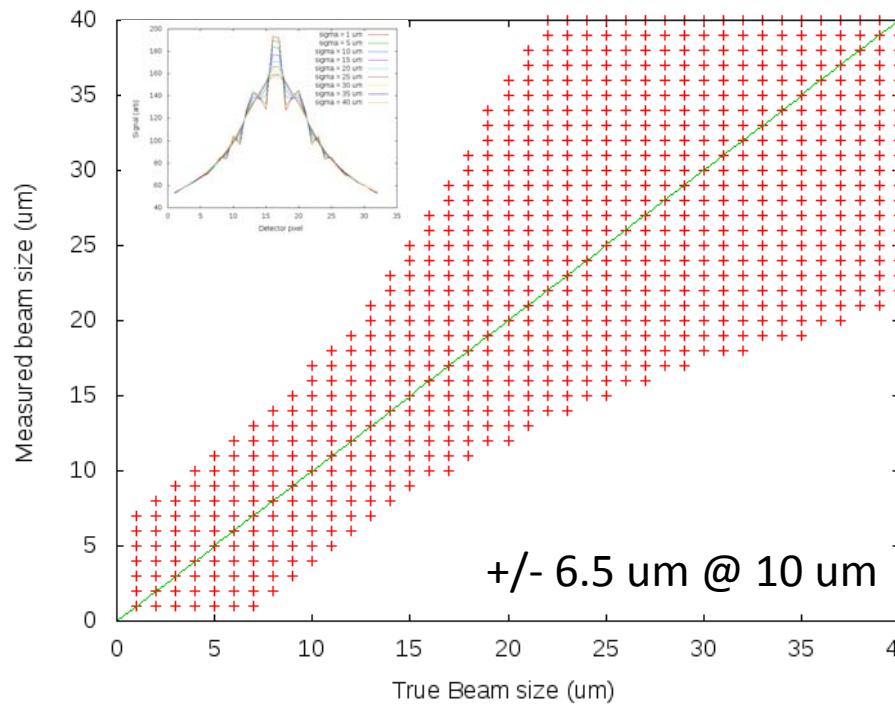
C Line, 5 μm 31-pixel CA
(Installed)



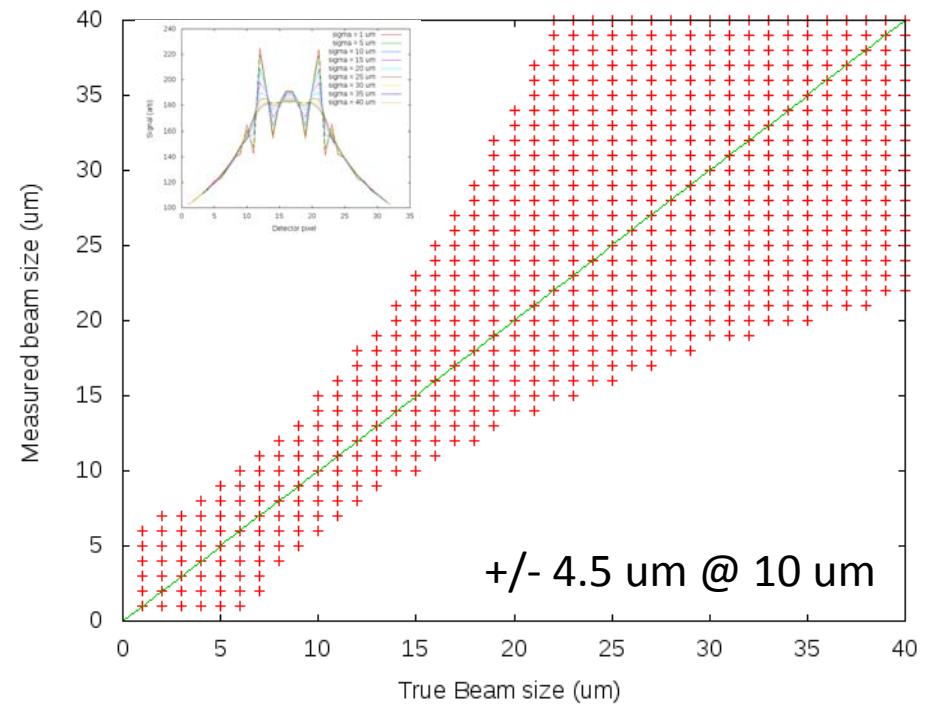
10 μm mask is better.

2 GeV FZP Resolutions

D Line, Applied Nanotools
(0.55 μm Au)



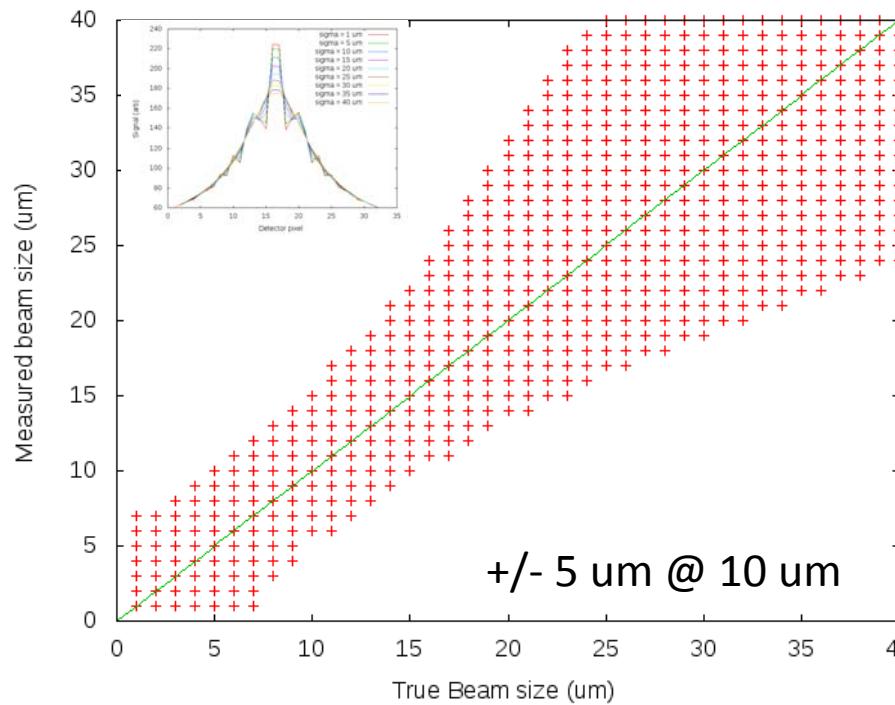
C Line, NTT-AT
(4 μm Ta)



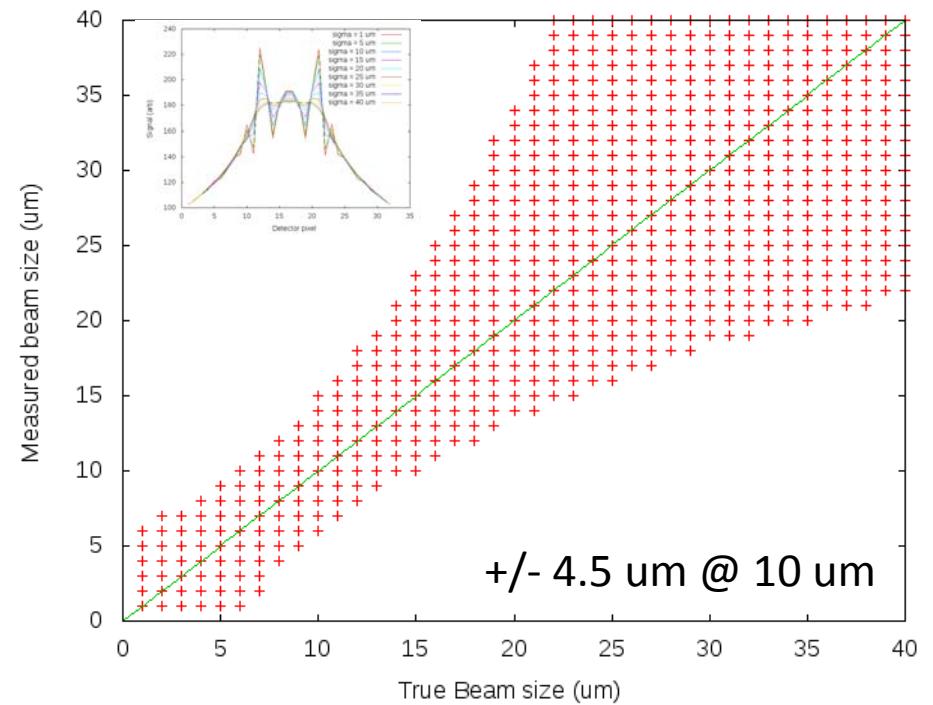
C Line slightly “better” due to mask material?
→ Better if template fitting, due to horns.
→ Worse if using FWHM.

2 GeV FZP Resolutions

C Line, Applied Nanotools
(0.55 μm Au)



C Line, NTT-AT FZP (installed)
(4 μm Ta)



Mask material does makes a difference in detector image,
though not so much in resolution achieved (if template fitting.)

Summary

- E-cloud data demonstrate ability to measure bunch-by-bunch, turn-by-turn size (down to ~ 10 um) and position data.
 - At larger beam sizes ($>\sim 50$ um), CA seems more prone to misfits than FZP.
- Method of estimating resolution seems to agree with CA data pretty well
 - Uncertainties in detector noise and inter-pixel calibration.
 - CA resolution: $\pm \sim 2$ um at 10 um.
 - FZP resolution: $\pm \sim 6$ um at 10 um.
- Resolutions estimated for all installed optical elements.
 - CA 10 um somewhat better than 5 um at both 2 GeV and 4 GeV.
 - Either CA better than FZP.
 - Applied Nanotools and NTT-AT FZPs show some feature differences due to mask material thickness. Which is better depends on what fitting method is used (template fitting or FWHM measurement.)
- We have elements installed now that can measure beam sizes to $<\sim 10$ um.
 - Can we go lower?
- We also have procedure that seems to work for evaluating designs in order to optimize resolution.

Spares

Xray Beam Size Monitors

- Coded Aperture Imaging:

- Technique developed by x-ray astronomers using a mask to modulate incoming light. Resulting image must be deconvolved through mask response (including diffraction and spectral width) to reconstruct object. Open aperture of 50% gives high flux throughput for bunch-by-bunch measurements. Heat-sensitive monochromator not needed.

Source distribution:

$$\begin{bmatrix} A_\sigma \\ A_\pi \end{bmatrix} = \frac{\sqrt{3}}{2\pi} \gamma \frac{\omega}{\omega_c} (1 + X^2) (-i) \begin{bmatrix} K_{2/3}(\eta) \\ \frac{iX}{\sqrt{1+X^2}} K_{1/3}(\eta) \end{bmatrix},$$

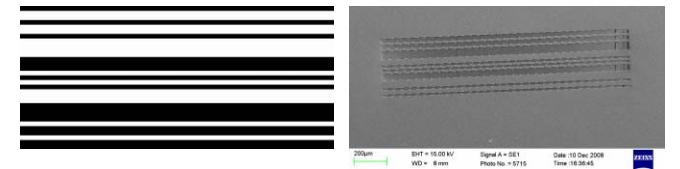
where

+

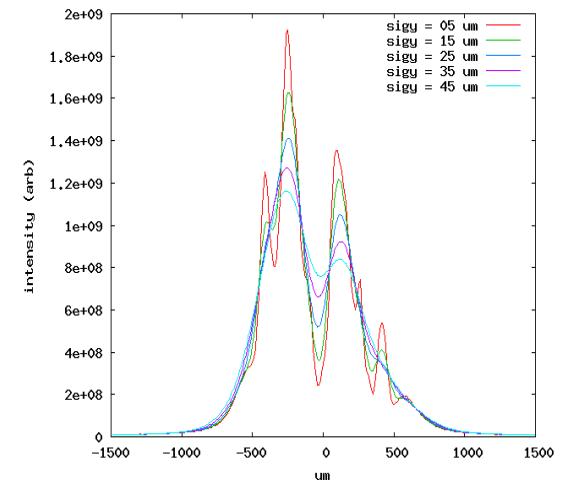
$$X = \gamma\psi,$$

Kirchhoff integral over mask
(+ detector response)
→ Detected pattern:

$$A_{\sigma,\pi}(y_d) = \frac{iA_{\sigma,\pi}(\text{source})}{\lambda} \int_{\text{mask}} \frac{t(y_m)}{r_1 r_2} e^{i\frac{2\pi}{\lambda}(r_1 + r_2)} \times \left(\frac{\cos \theta_1 + \cos \theta_2}{2} \right) dy_m,$$



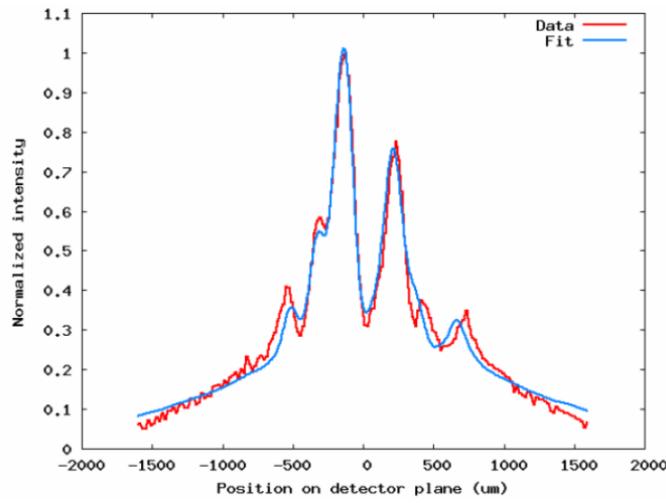
Uniformly Redundant Array (URA) for x-ray imaging being tested at CesrTA



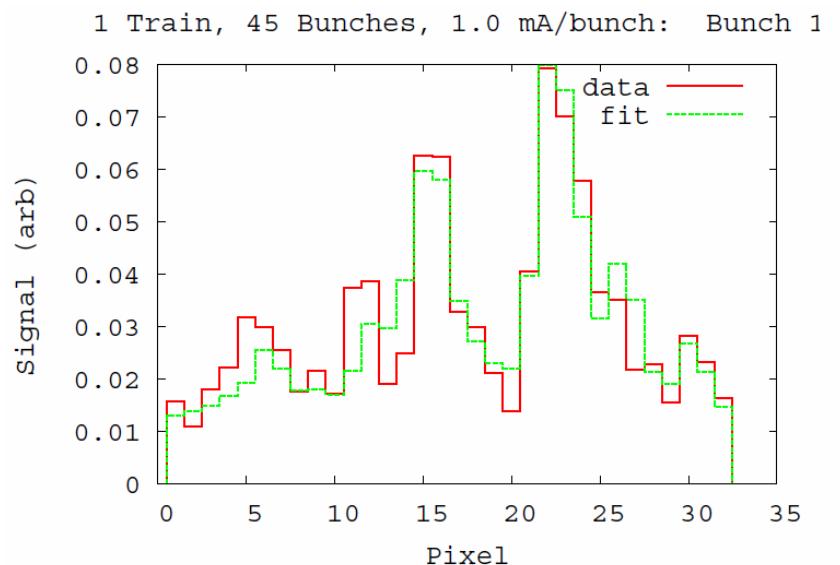
Simulated detector response for various beam sizes at CesrTA

Data Analysis

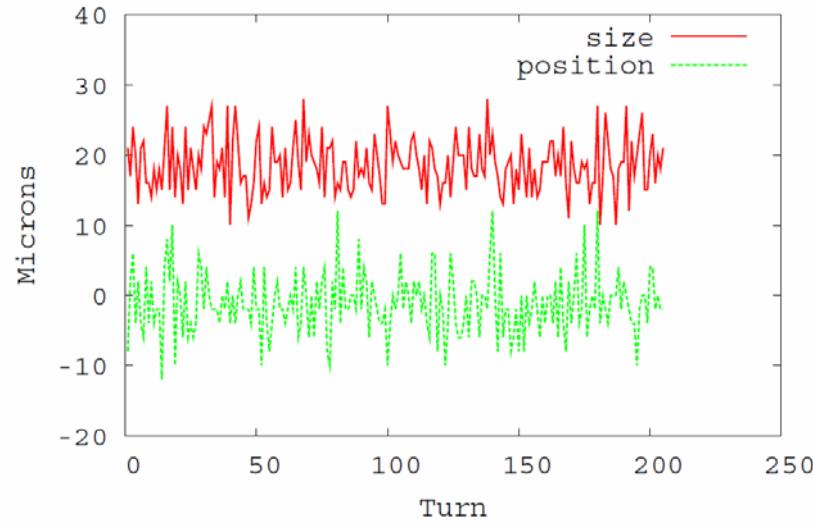
- 1) Simulate point response functions (PRFs) from various source positions to detector, taking into account beam spectrum, attenuations and phase shifts of mask and beamline materials, and detector response.
- 2) Add PRFs, weighted to possible proposed beam distributions.
- 3) Find best fit to detector data.



Measured slow-scan detector image (red) at CesrTA, used to validate simulation (blue)

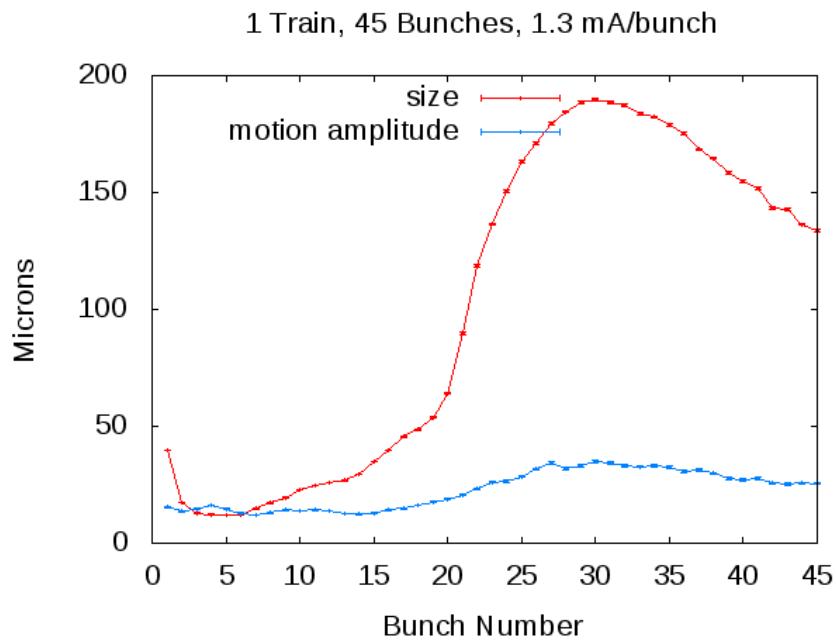


Example of single-shot data (single-bunch, single-turn)



Example of turn-by-turn data (one bunch out of train)

Example of electron-cloud study data taken at CesrTA



Bunch-by-bunch size and centroid RMS motion.
4096 turns, fitted separately then averaged for each bunch.
(Error bars statistical.)

Fourier power spectrum of fitted beam position

Progress so far at CesrTA:

- Validated simulation code with real beam.
- Demonstrated measurement of beam sizes down to $\sim 10 \mu\text{m}$ (preliminary).
- Demonstrated bunch-by-bunch, turn-by-turn size measurements.

