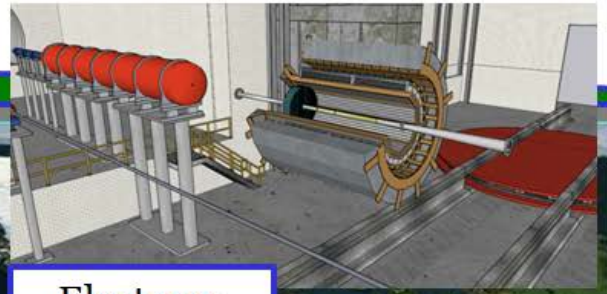
University of Hawaii
Instrumentation Development Laboratory

IDL Oveview
(last picture/lab party before the pandemic)
24-JAN-2023

IDL
Instrumentation
Development
Lab

A world-wide impact

Deep Underground Neutrino Experiment (DUNE)



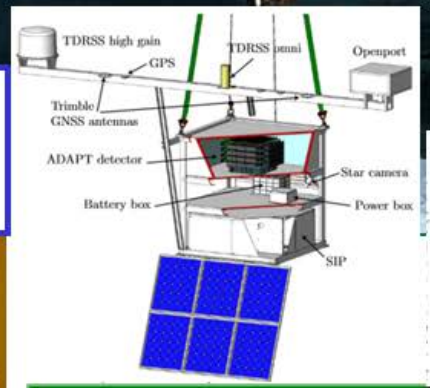
Electron-Ion Collider

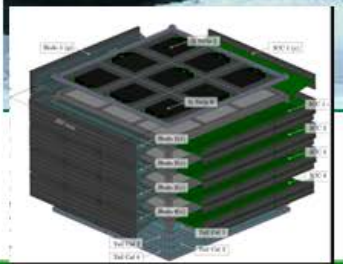
Belle II Construction & Commissioning (pixel, Silicon Upgrades)

Belle II

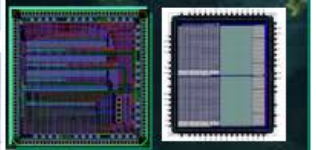
Cherenkov Telescope Array (CTA)

APT prototype (ADAPT)



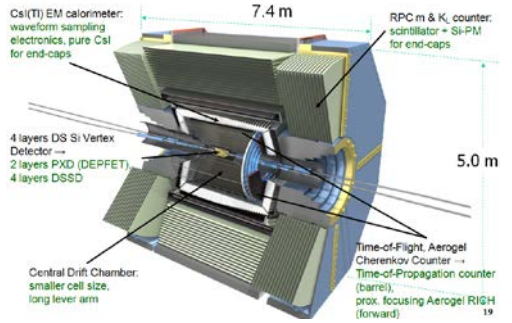


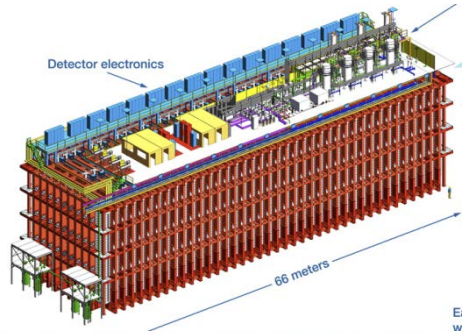
ANITA 1-4 PUEO

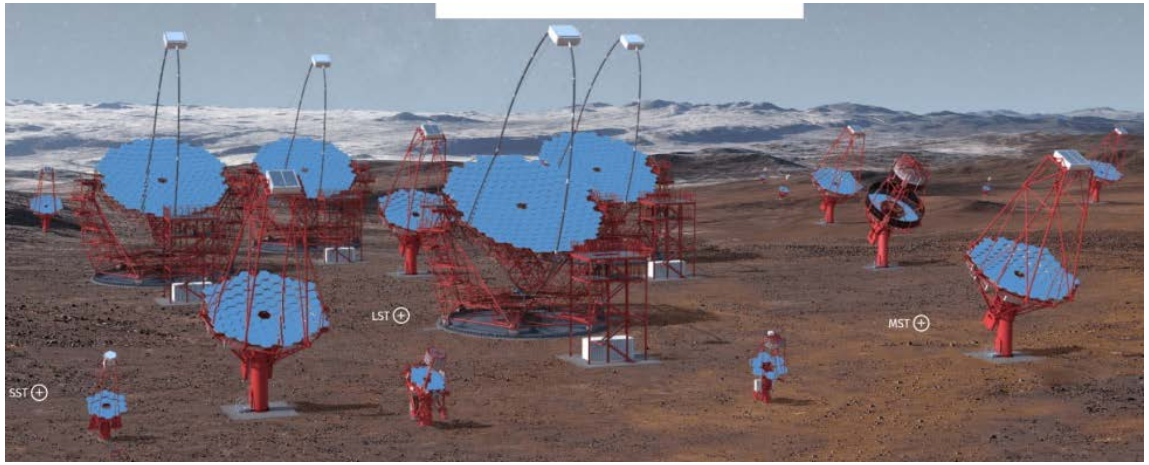


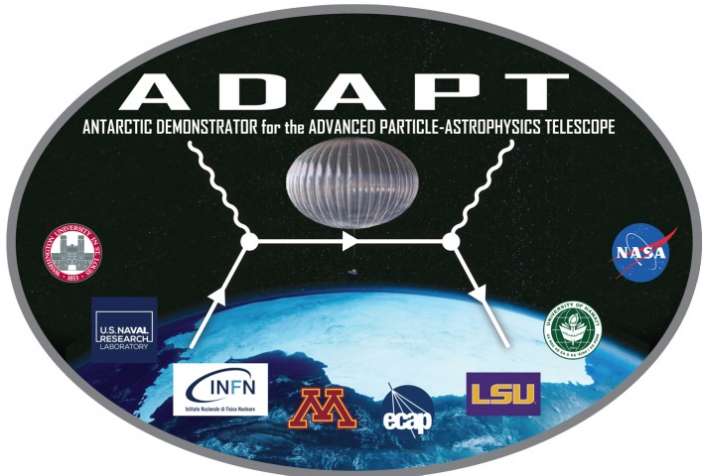
ASICs as enabling tools to the Big Questions

1. CP violation (how get more matter than anti-matter, but not too much)
2. Does CP violation also occur in leptons (and 3-generation problem) ?
3. Origin of highest energy cosmic rays and their neutrino fingerprint?
4. What is the nature and origin of Dark Matter ?



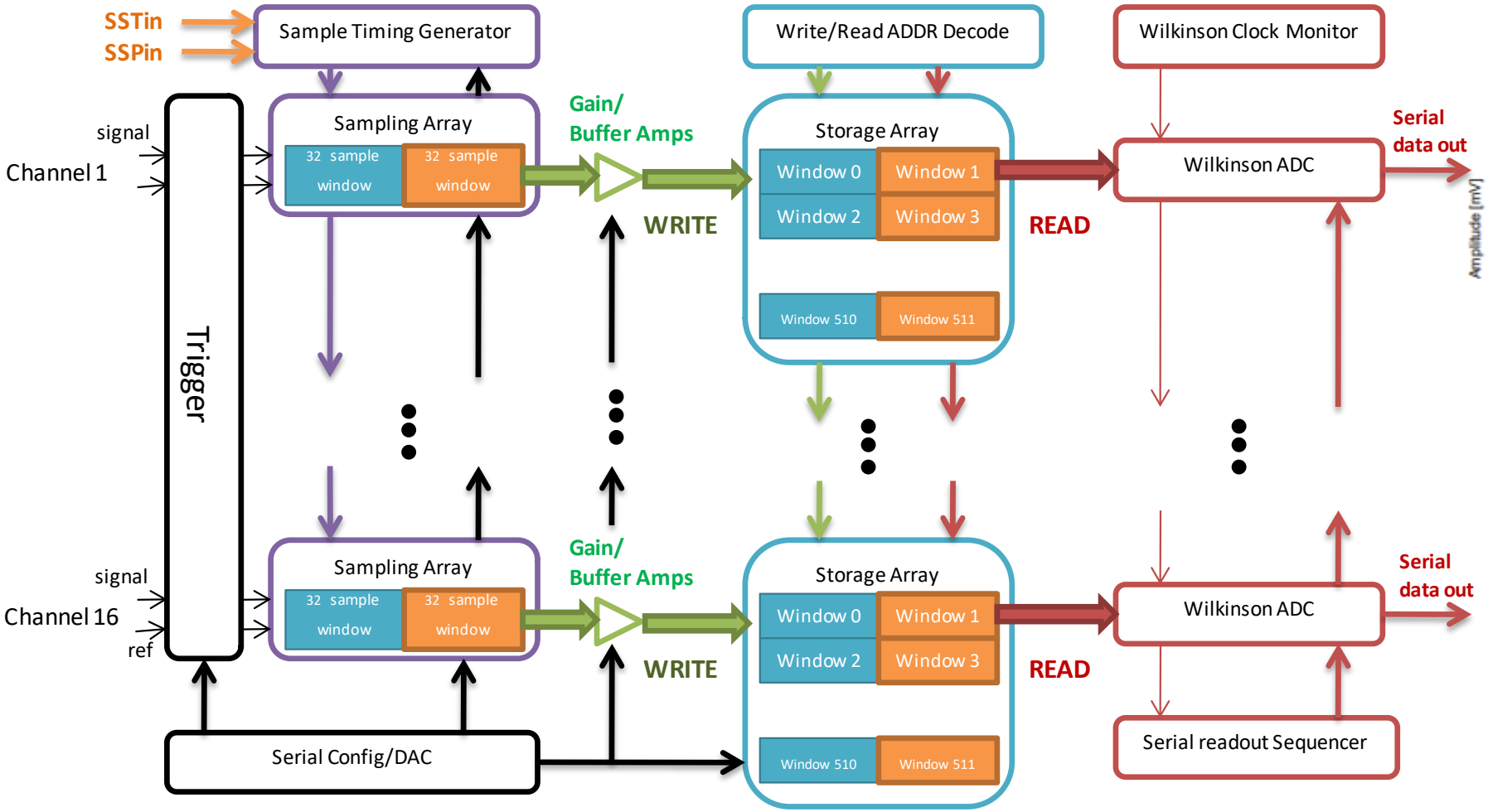


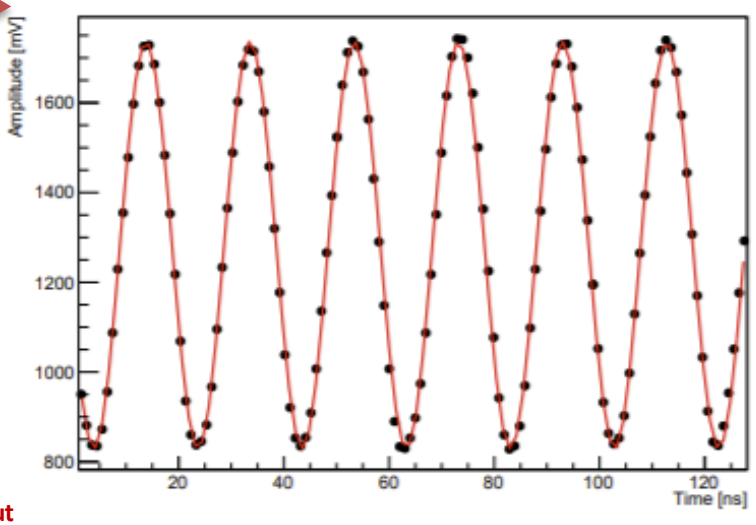




Enabling technology: GigaSample/s, low power

- **Transient events, high channel density**

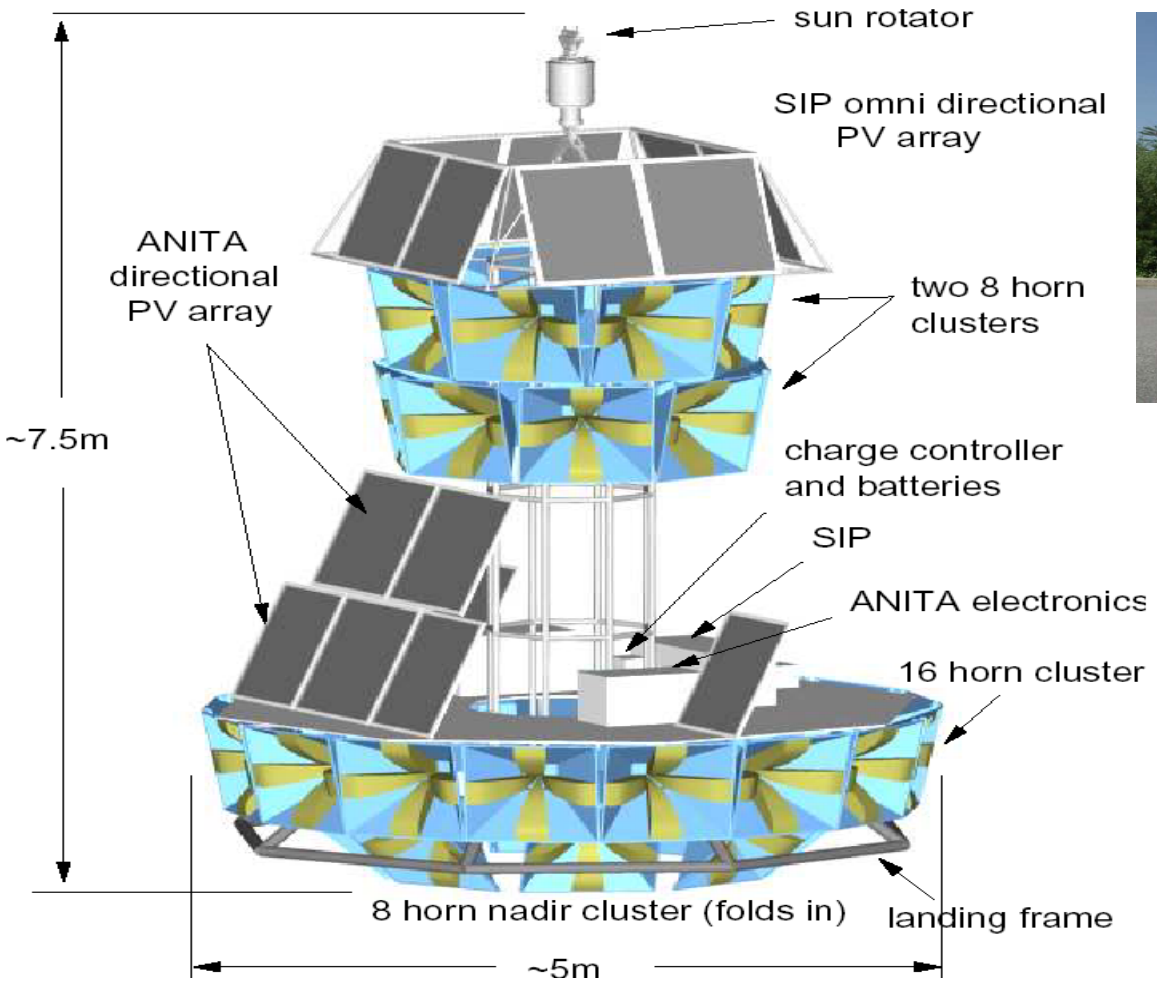


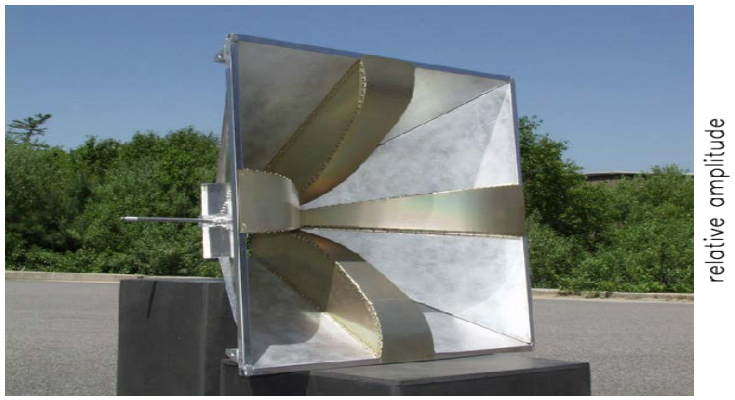


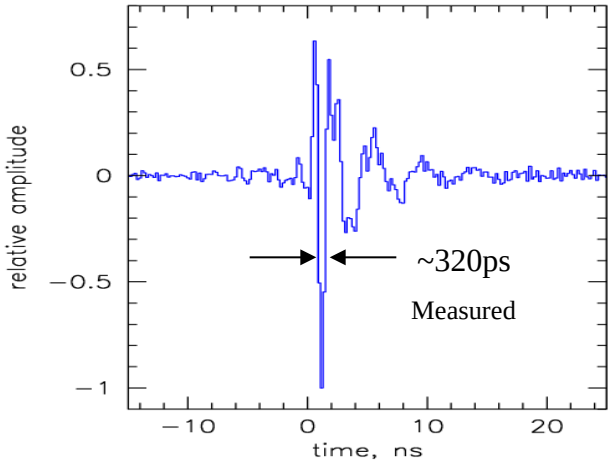
- Sample 50MHz sine wave used for calibration

A Very Challenging Design

A radio "feedhorn array" for the Antarctica Continent







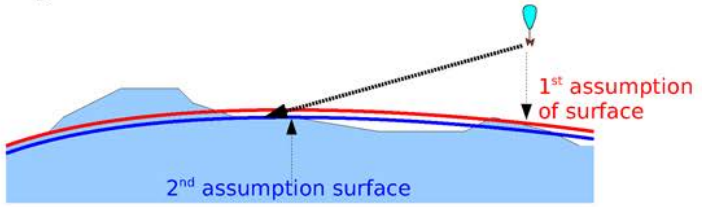
- Quad-ridged horn antennas provide superb impulse response & bandwidth (200-1200 MHz)
- Interferometry & beam gradiometry from multiple overlapped antenna measurements

ANITA: LABRADOR3

- **Flown all 4 ANITA flights. LAB4D developed for ANITA5**

<30ps timing

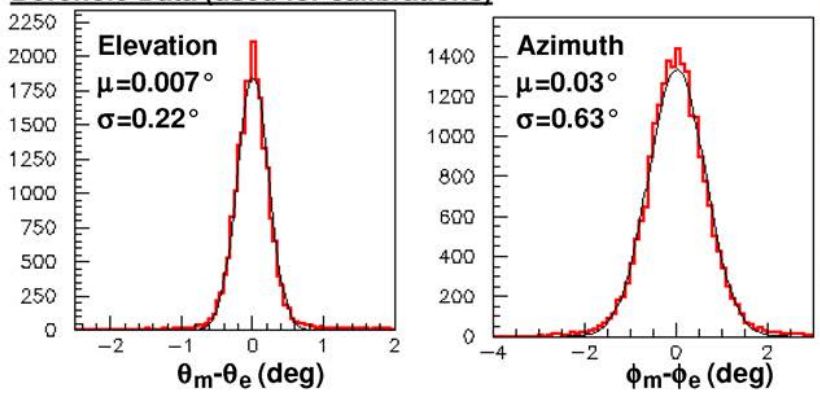
RF Projection onto the surface



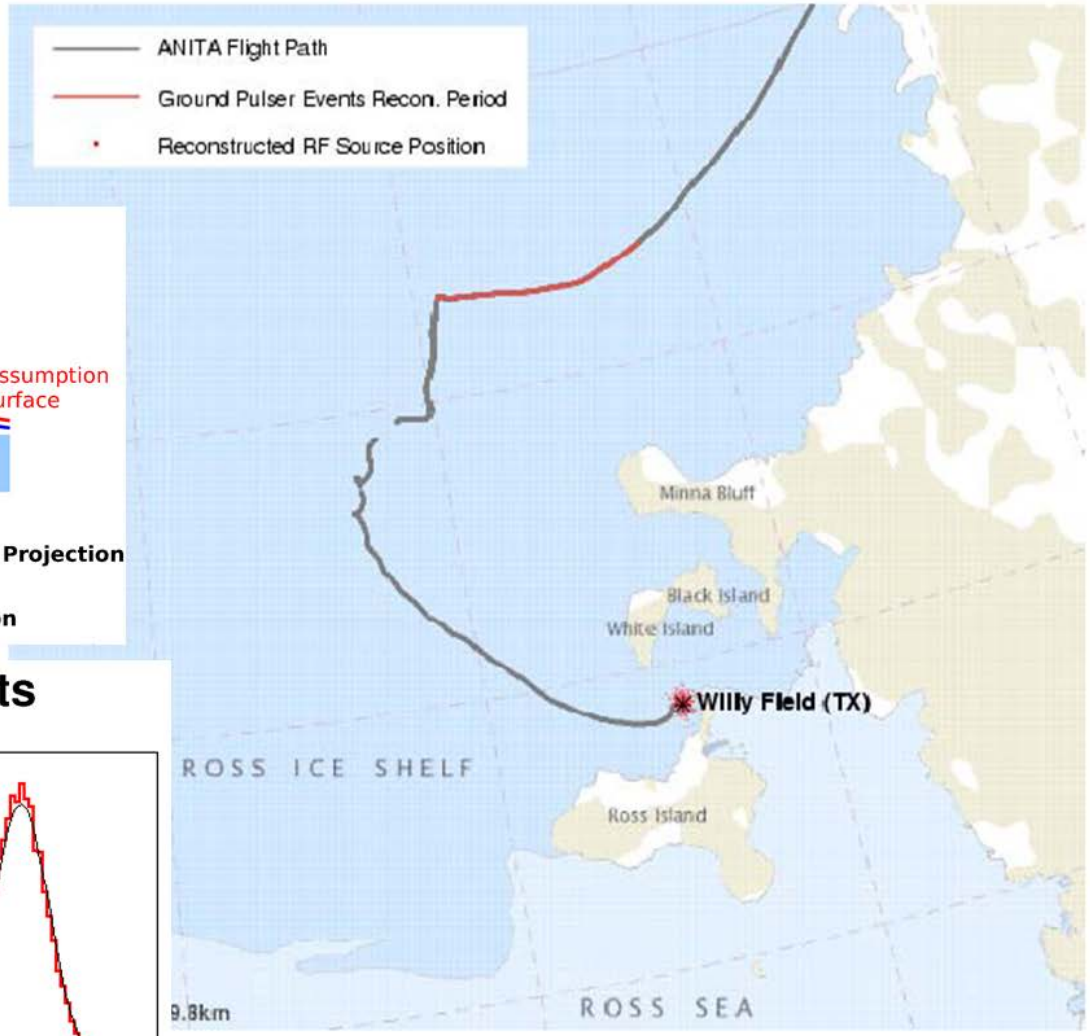
Fast Algorithm: Line Sphere intersection
1st R_{earth}=Geoid + Surface @ Ballon position -> Rough Projection
2nd R_{earth} = Geoid + Surface @ (position from 1st)
3rd: one more iteration -> converged after 2nd iteration

V-pol results

Borehole Data (used for calibrations)

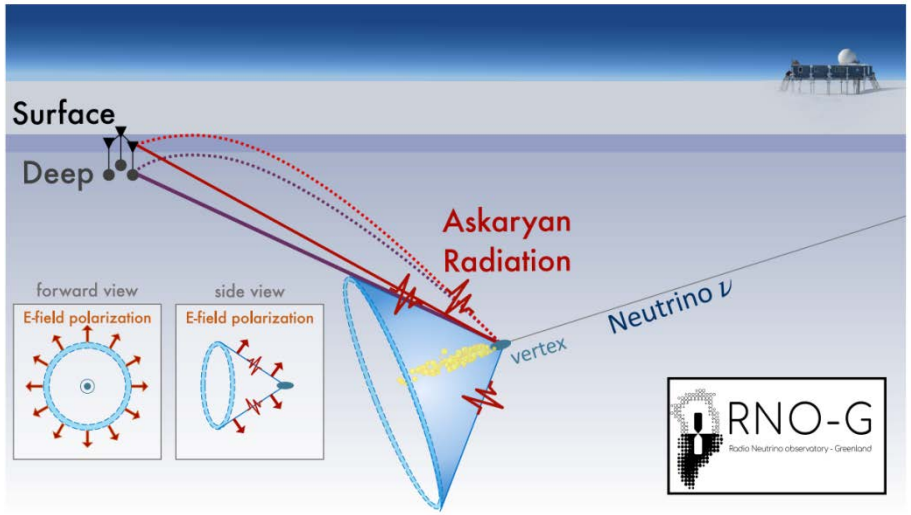


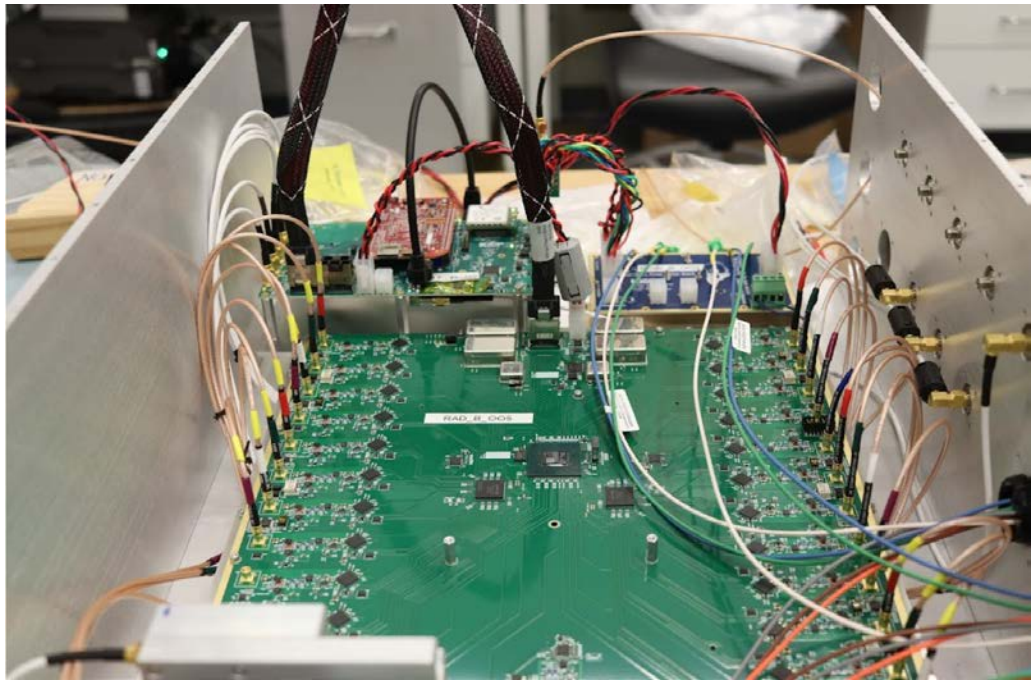
3.8 mrad **11.2 mrad**

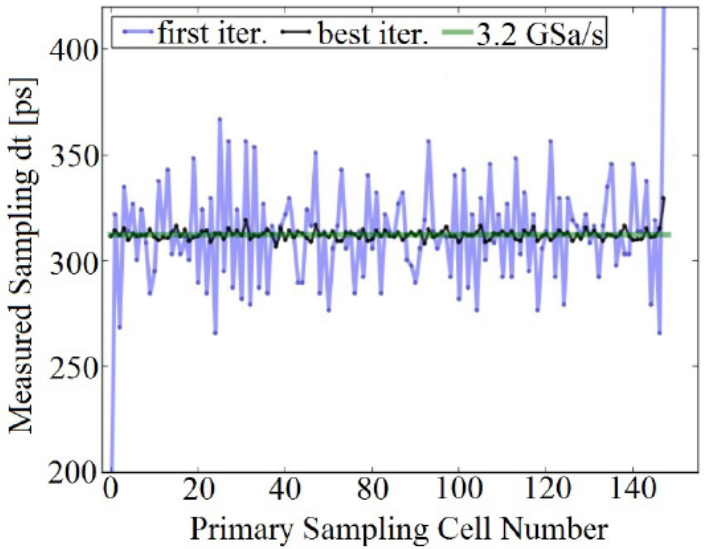


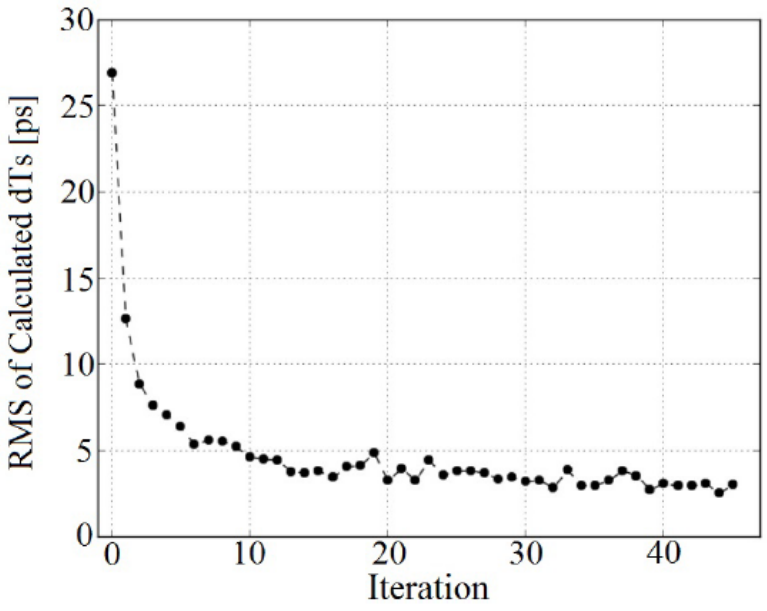
RNO-G: LAB4D

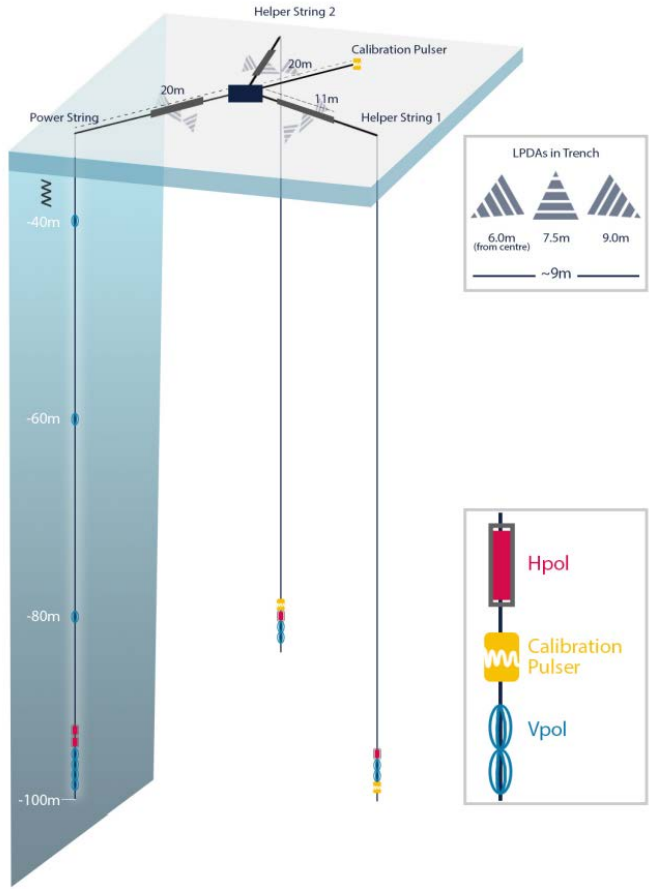
- **LAB4D explored removing "dT" non-uniformity**







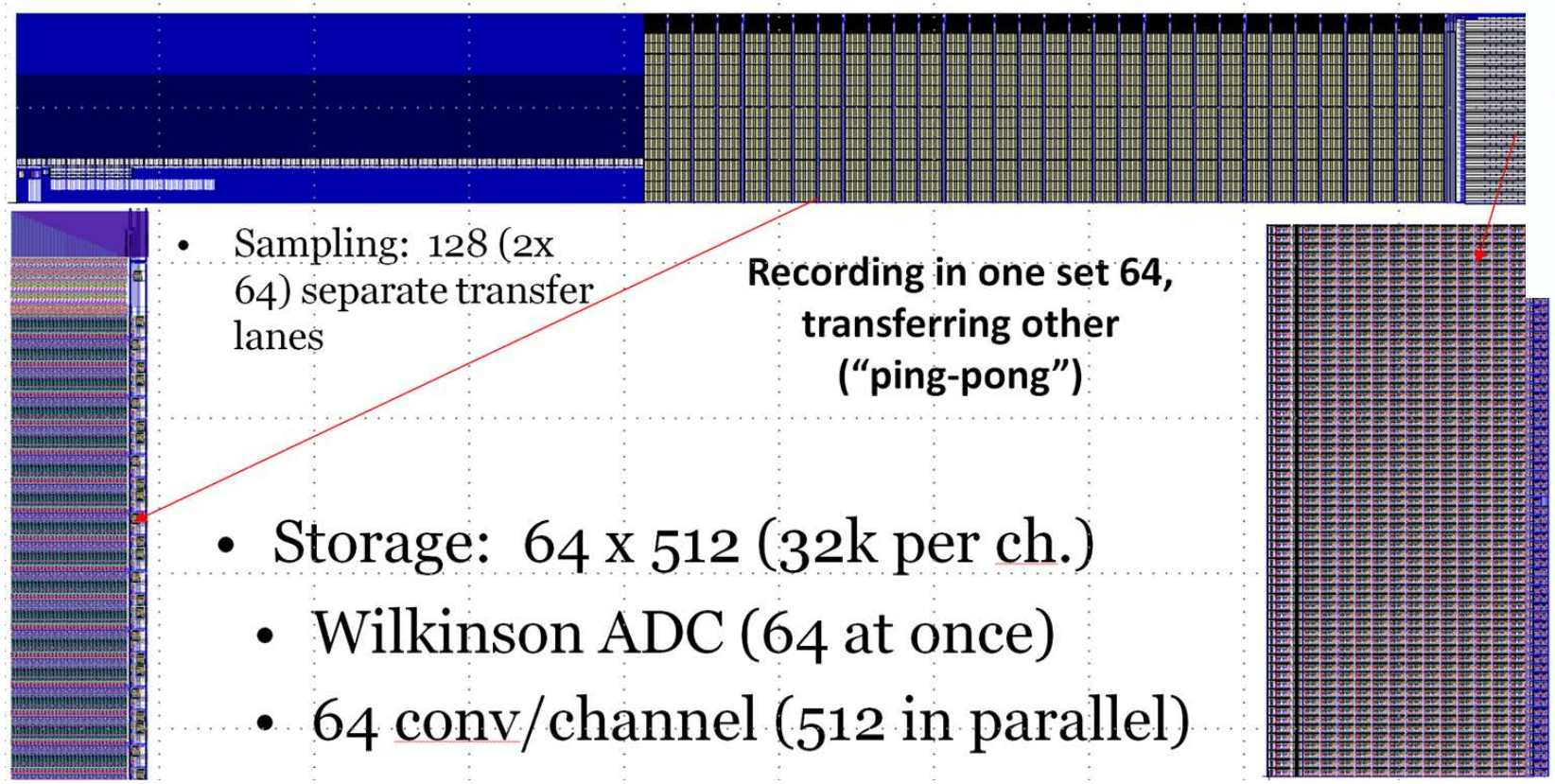


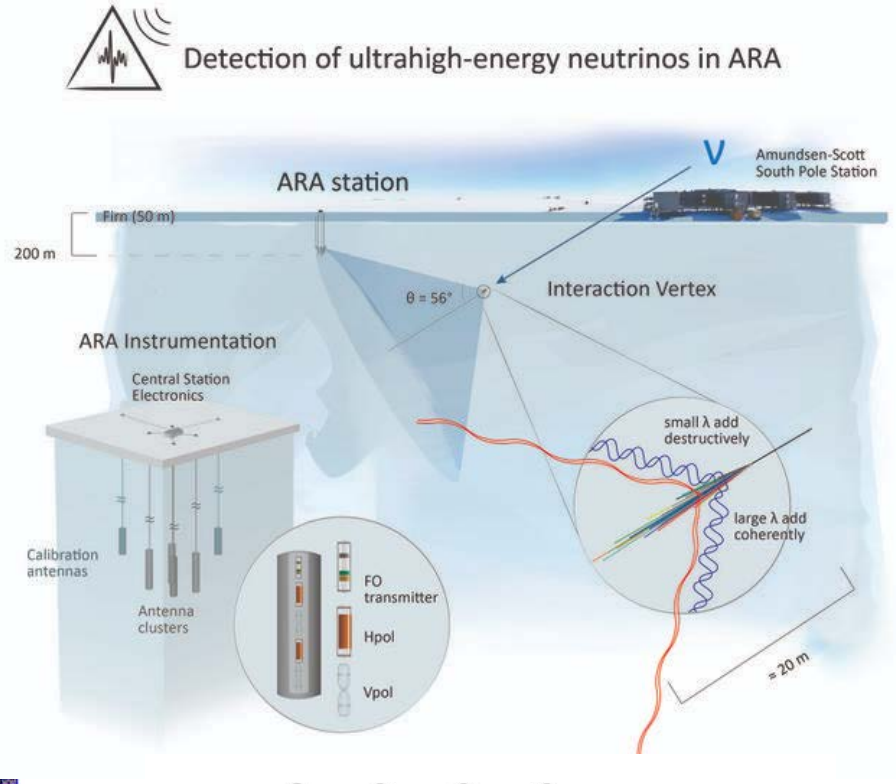


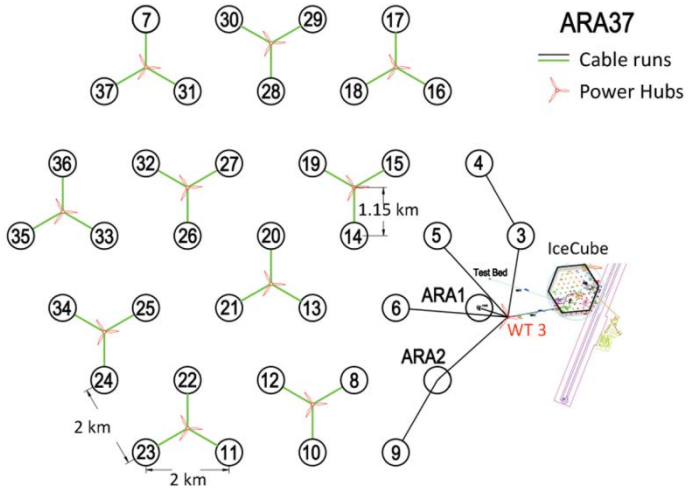
- Removes ~15% diff Timing sample width
- Cal. Converges to ~3ps in 32 iterations
- Avoids re-processing and splining (impossible for real-time triggering)

ARA: IceRadio Sampler

- **Explore how deep (reasonably) can go with SCA architecture**
- **Force single-station trigger and coherently phase stations offline**







Belle II Detector

BEAST (Background commissioning detector)

KLong and muon detector:
Resistive Plate Chambers (barrel outer layers)
Scintillator + WLSF + SiPM's (end-caps , inner 2 barrel layers)

EM Calorimeter:
CsI(Tl), waveform sampling (barrel+ endcap)

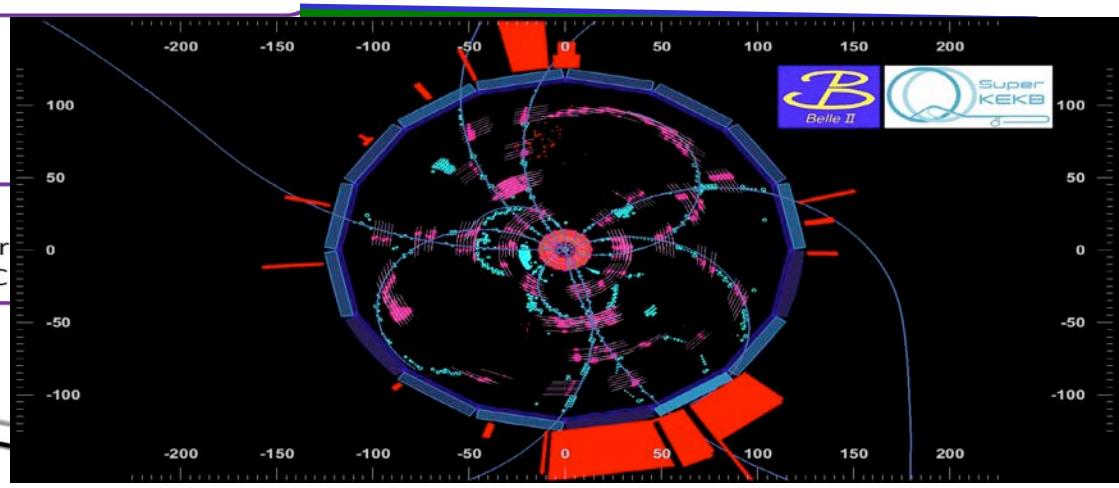
Particle Identification
TOP detector system (barr
Prox. focusing Aerogel RIC

electrons (7 GeV)

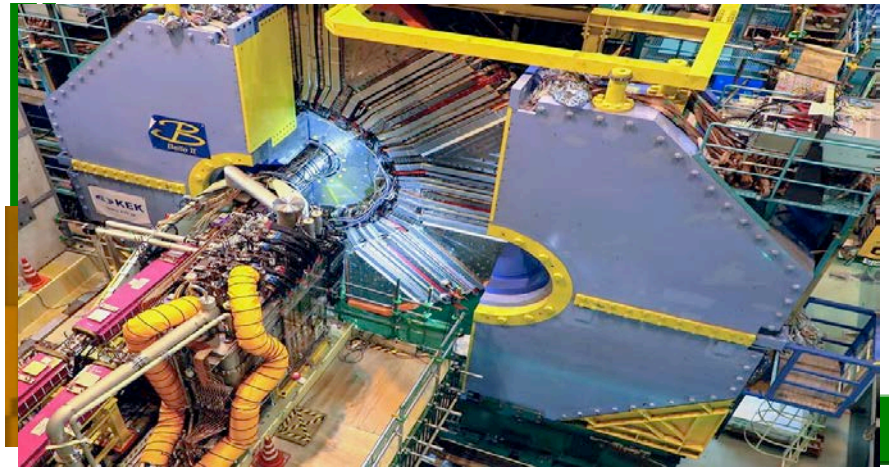
Beryllium beam pipe
2cm diameter

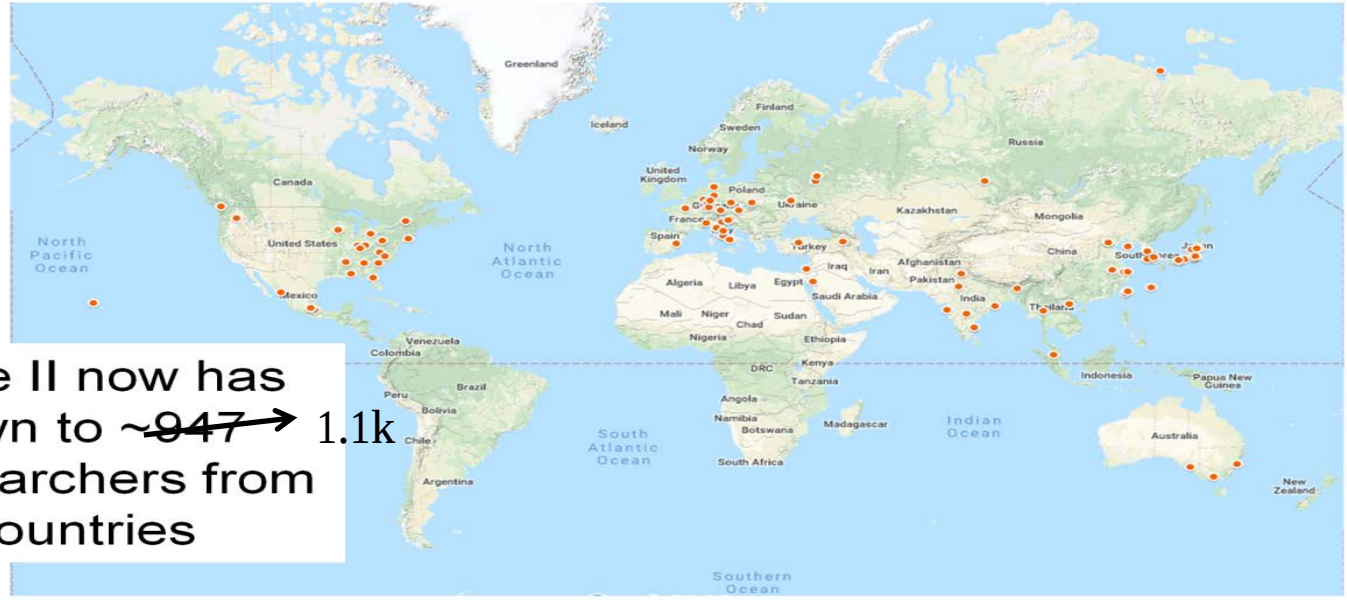
Vertex Detector
2 layers DEPFET + 4 layers DSSD

Central Drift Chamber
He(50%):C_2H_6(50%), small cells, long lever arm, fast electronics (Core element)



positrons (4 GeV)



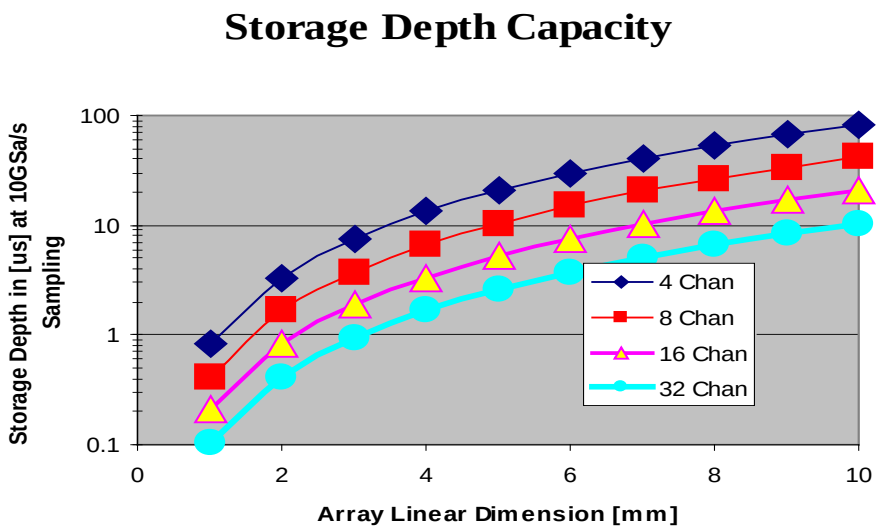


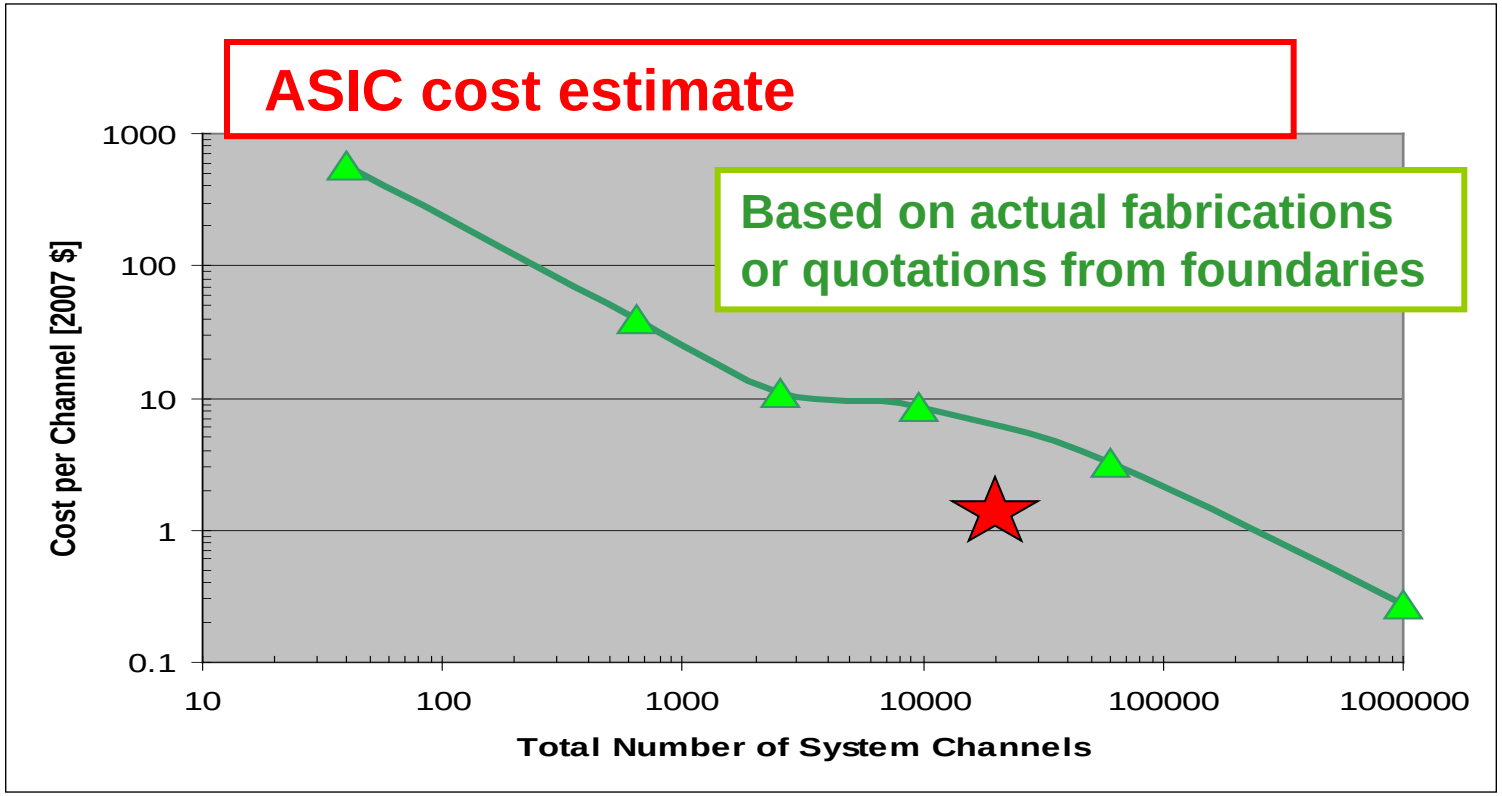
Belle II now has grown to ~~947~~ 1.1k researchers from 26 countries

Looking back on >10 year development

- ASIC costing well understood, very competitive!

NIM A591 (2008) 534-345.

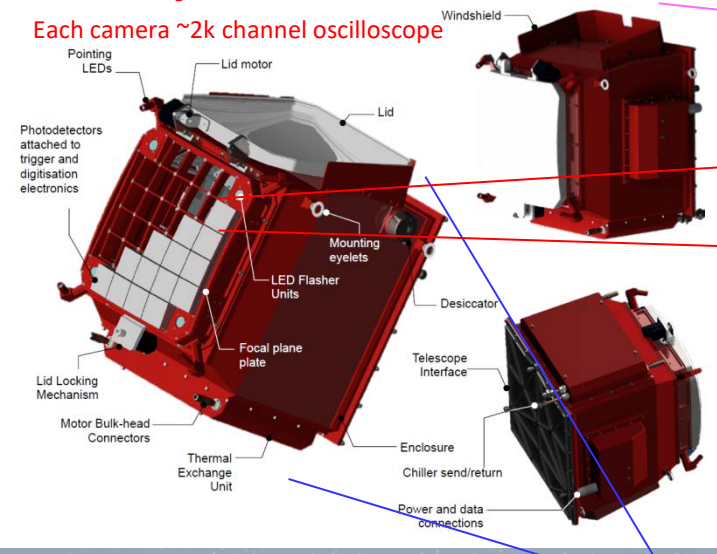


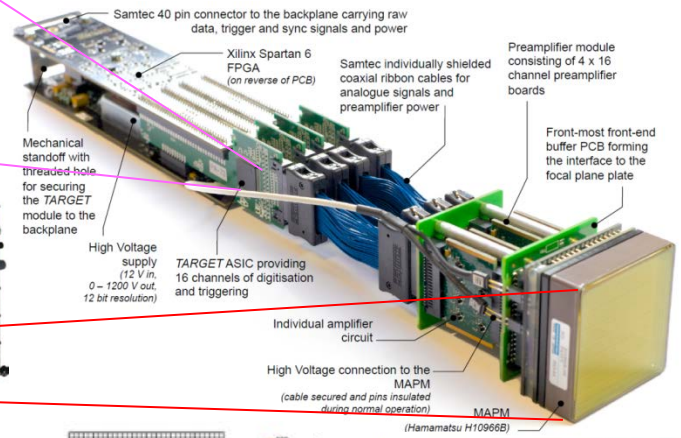


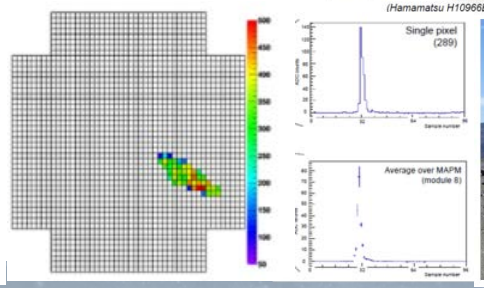
Cherenkov Telescope Array

- **Enabling in reducing the SST camera production costs**

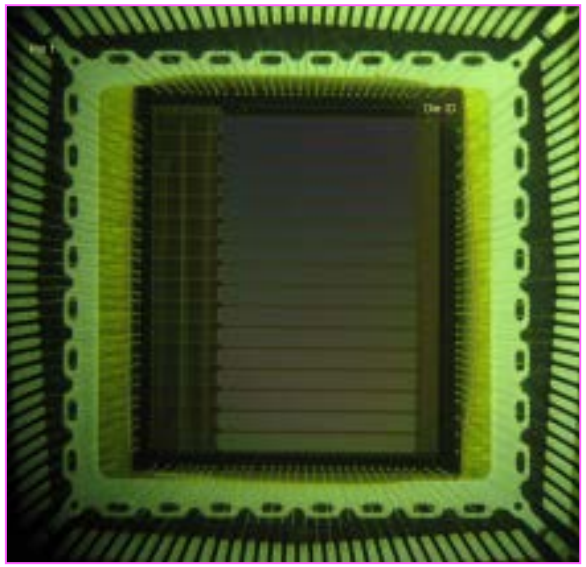
Each camera ~2k channel oscilloscope



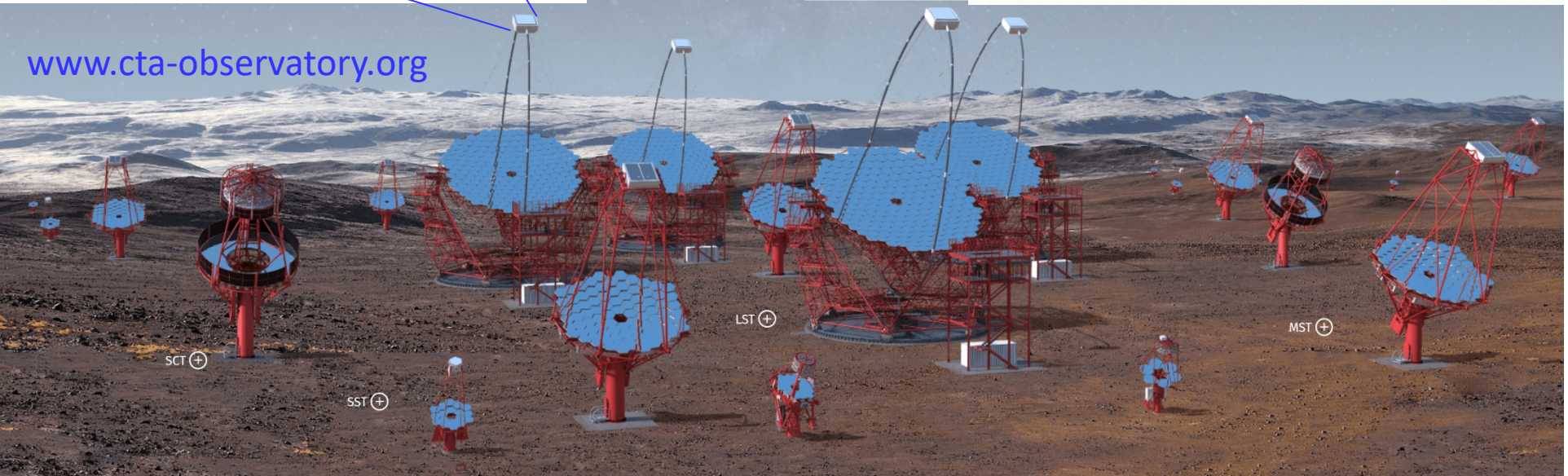




CTC (sampler)
CT5TEA (trigger)
Gang Liu

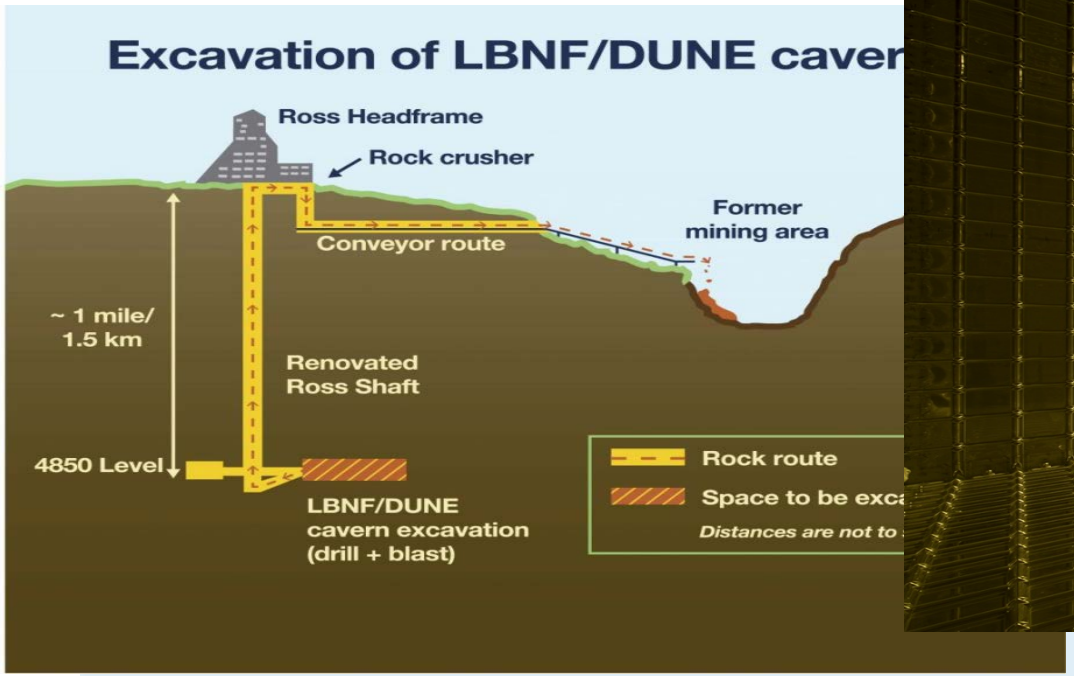


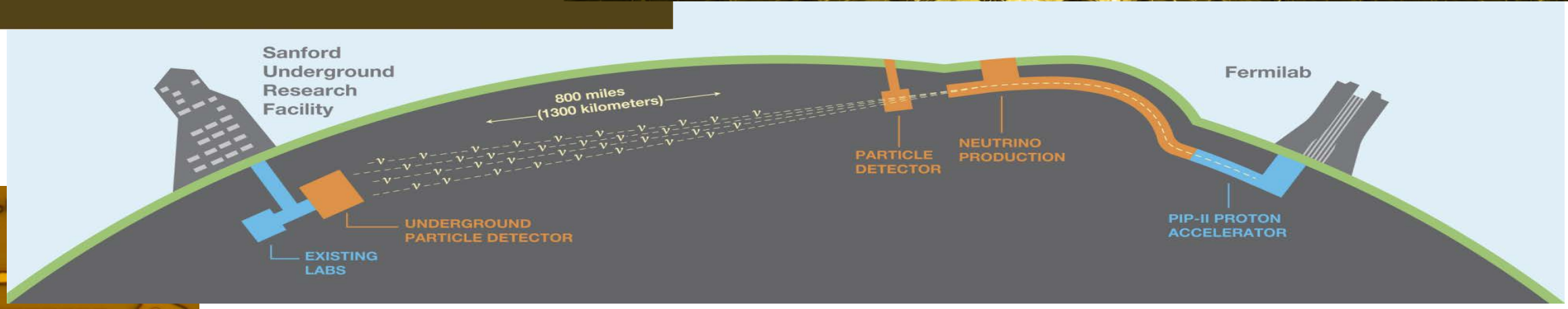
www.cta-observatory.org



- First production batch of 6k ASICs (~100k chan) started
- ALPHA configuration baseline approved, construction started
- 2023 first light (2026 complete) – though eruption in north

Deep Underground Neutrino Experiment





APT: ALPHA

- **Lower power, higher density readout (ganging readout/FPGA)**

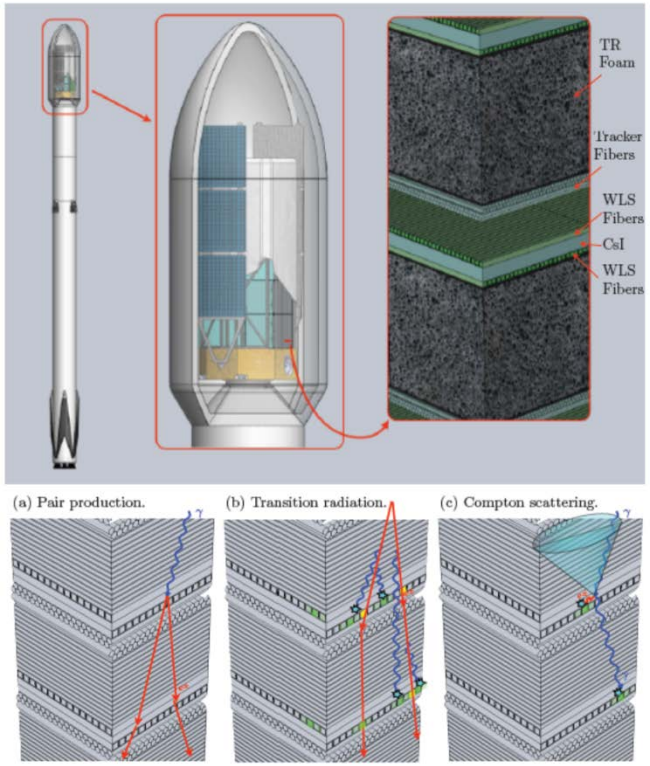
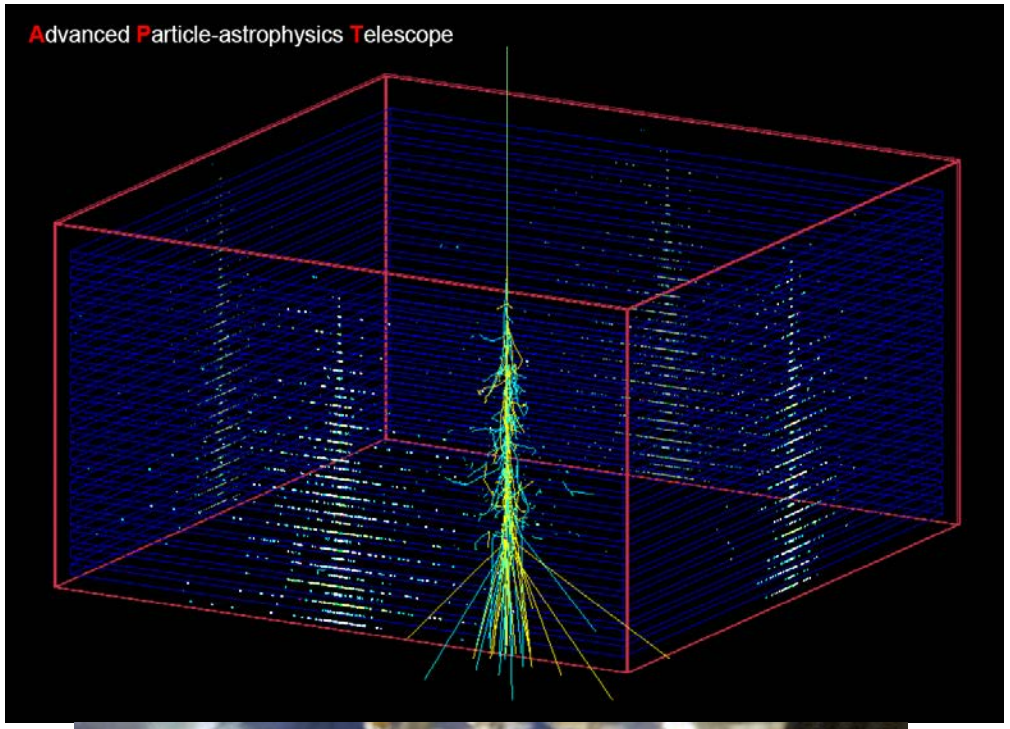
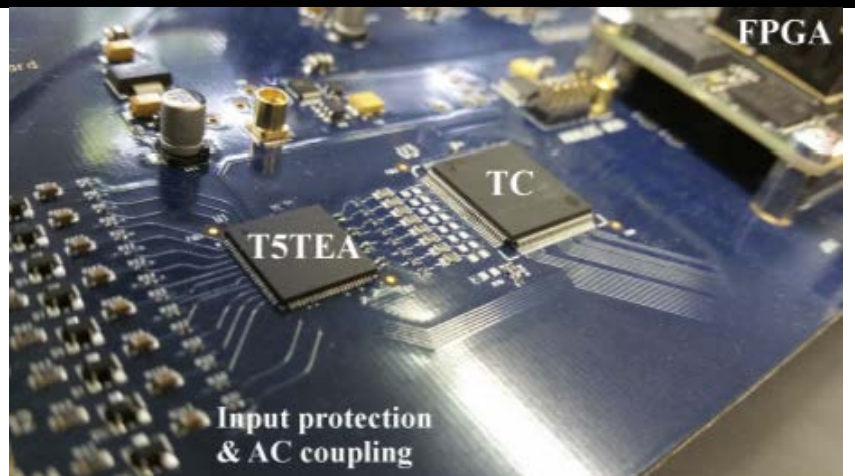


Figure 1: *Top: APT in Falcon-9 faring. Bottom: APT detection modes.*





Baseline design: 380,000 channels (needs to be low power – happy with TARGETC + T5TEA functionality)

- **Triggering seems fine**
- **"hang" more digitizers from a single FPGA (data is sparse)**
- **No central front-end trigger**

Discussion

- **Gap between bleeding-edge technology nodes used in particle and astroparticle physics is increasing**
- **There are strong economic reasons for this**
- **One impact is that it is increasingly difficult to attract students that want to work in the hottest/newest technology nodes**
- **Beyond affording fabrication costs, in recent years mere access to these processes has been increasing difficult/impossible due to breakdown of the NDA process**

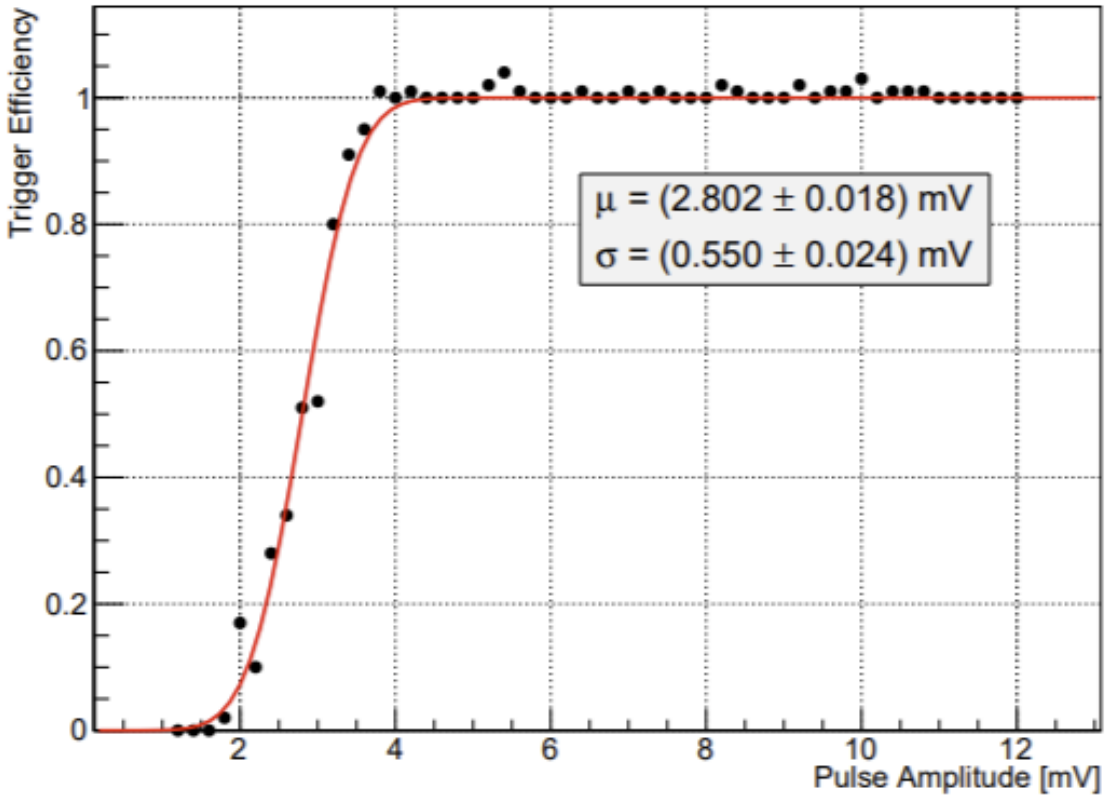
Backup Slides

Further exploration directions?

- **Limits of timing resolution (RF-Pix → <1ps time resolution)**
- **Channel density**
- **Processing/feature extraction on chip (digital synthesis)**
- **Timing calibration**

CTA: CT5TEA

- **Low threshold, analog sum triggering (companion ASIC)**

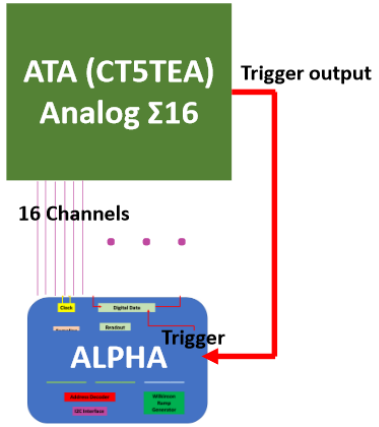


- **Sum of 4 inputs achieves this**
- **Provides DC offset to TARGET sampler**

ALPHA Readout Concept

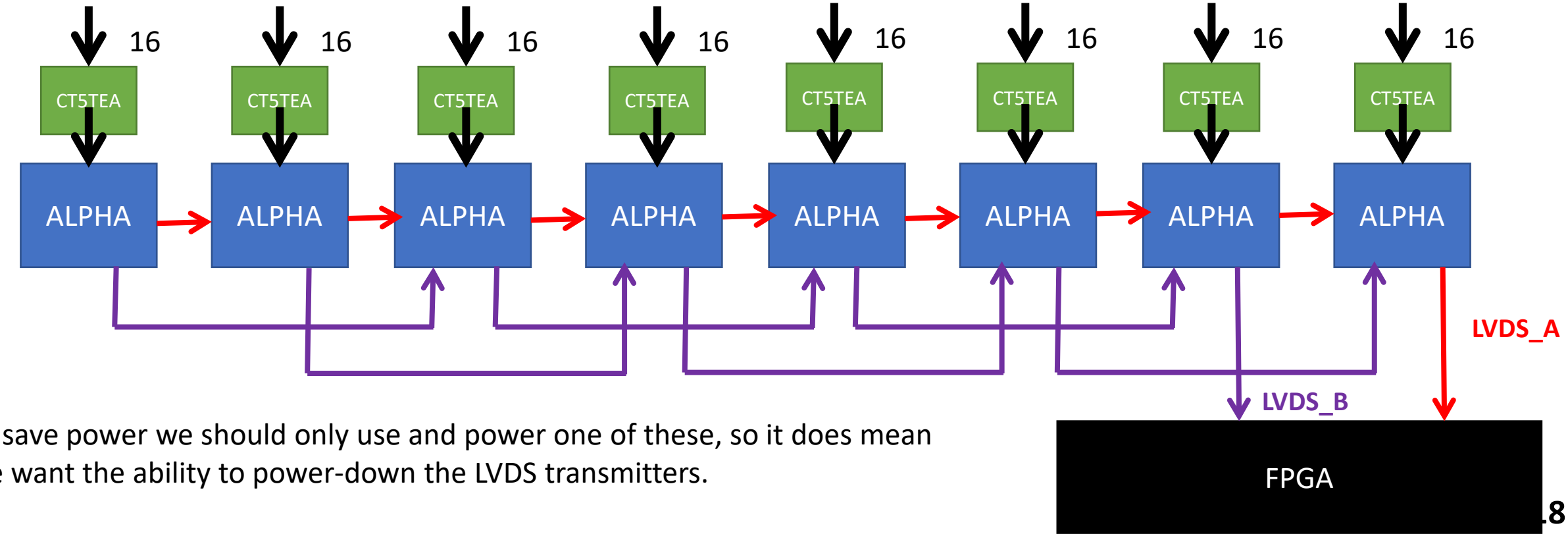
Use existing Cherenkov Telescope Array trigger ASIC (tweaked to have an analog sum-of-16 trigger) and operate digitization of each ALPHA independently.

Maximize the number of ALPHA served by a single FPGA (power, I/O limits)
In theory, 64 chains * 32 ALPHA * 16 channels ~ 2^15 channels/FPGA



ATA
1. Triggering
2. Vped offset

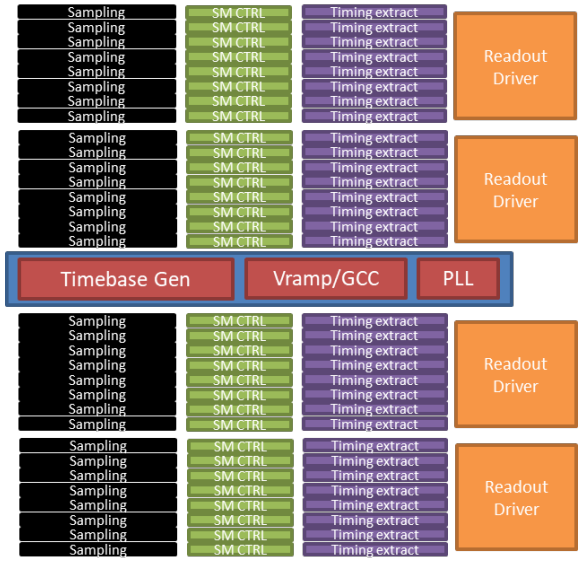
One ATA for one ALPHA

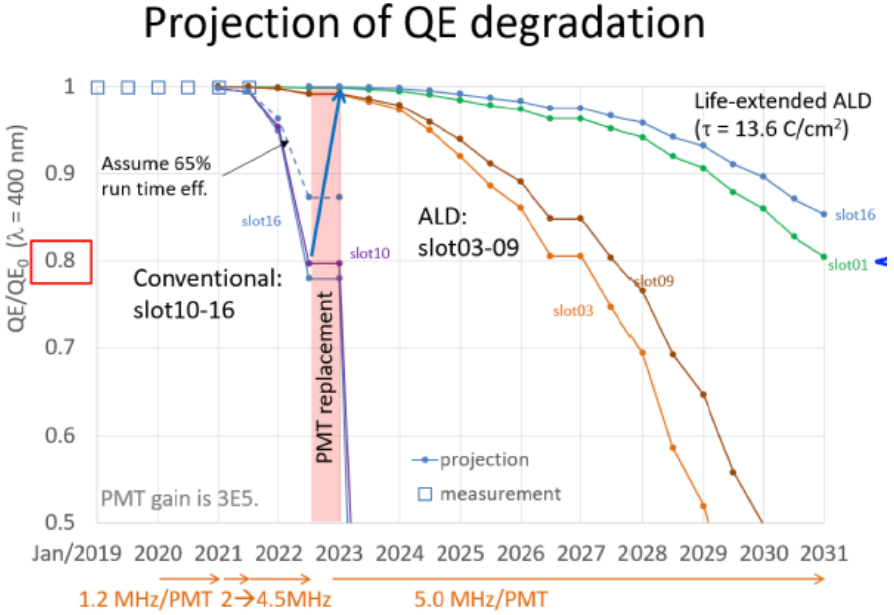


To save power we should only use and power one of these, so it does mean we want the ability to power-down the LVDS transmitters.

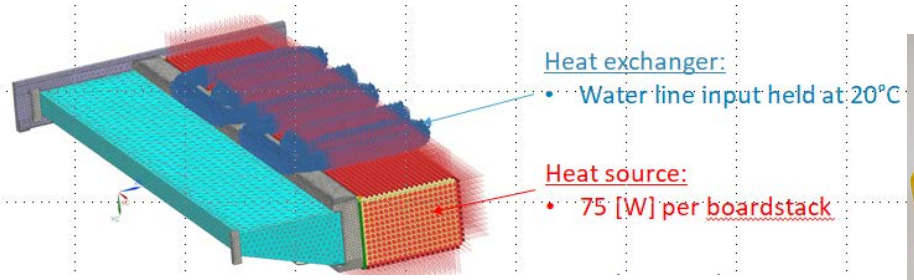
GULFstream Overview (iTOP motivation)

- **Much more compact, lower power**
- **No long storage depth – digitize and ship all hits**
- **GULFstream Features**
  - 32 channels
  - 4x deep, 64 sample SCA
  - 4x Gbps data links (8ch/link)
  - 32 bit words (3 chan, coarse/fine time)
  - Hit streaming: 3.2 Mhits/s/chan
  - Each channel operates independently
  - ~100ns/6-bit conversion
  - Direct feature extraction
  - No pedestal acquisition
  - To user, looks like a TDC
  - Low power/channel (exact savings to be confirmed)

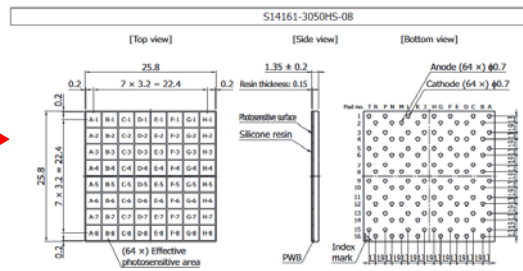




2.6M$ to replace MCP-PMTs (and 4-5 years!)



Operating @ thermal redline

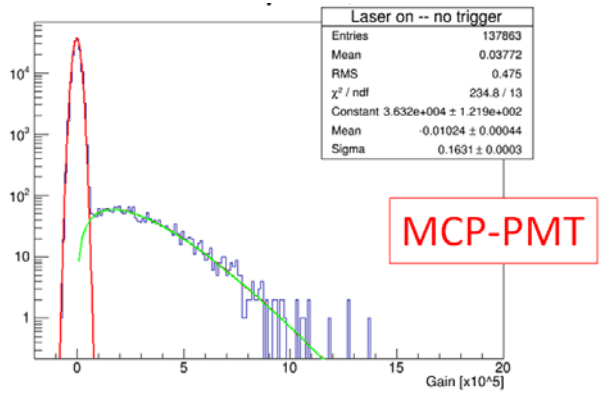


4x more readout channels!

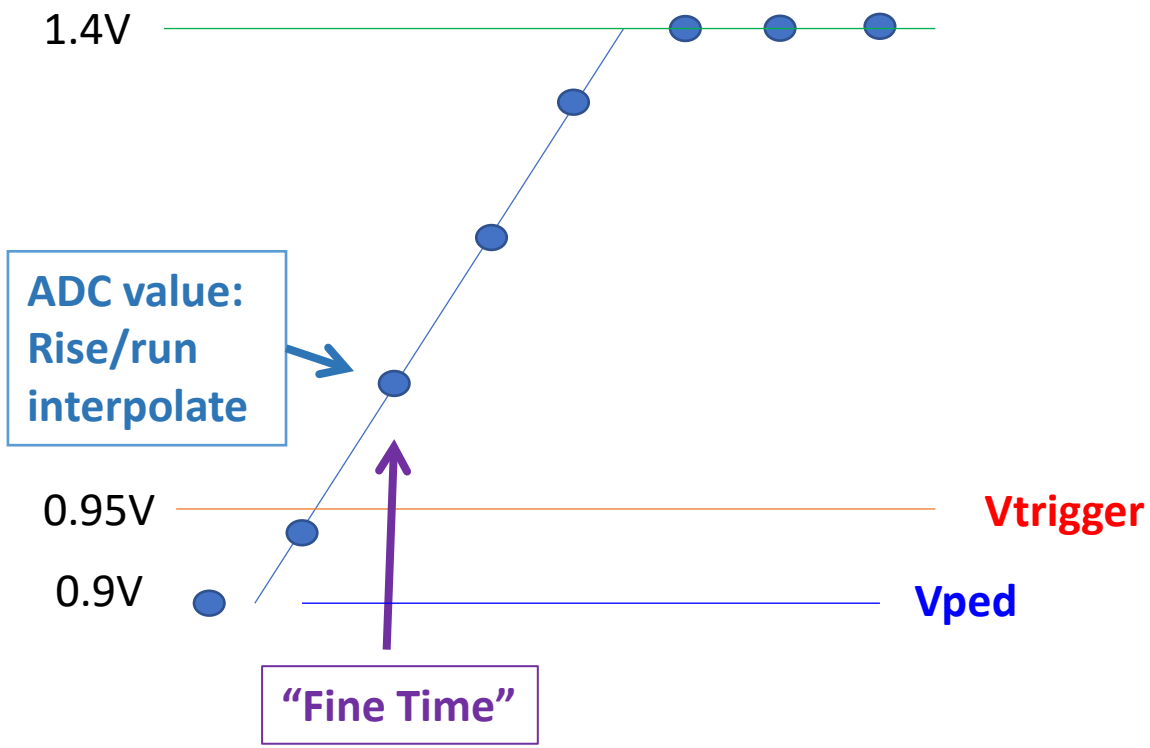
-30C SiPM operation

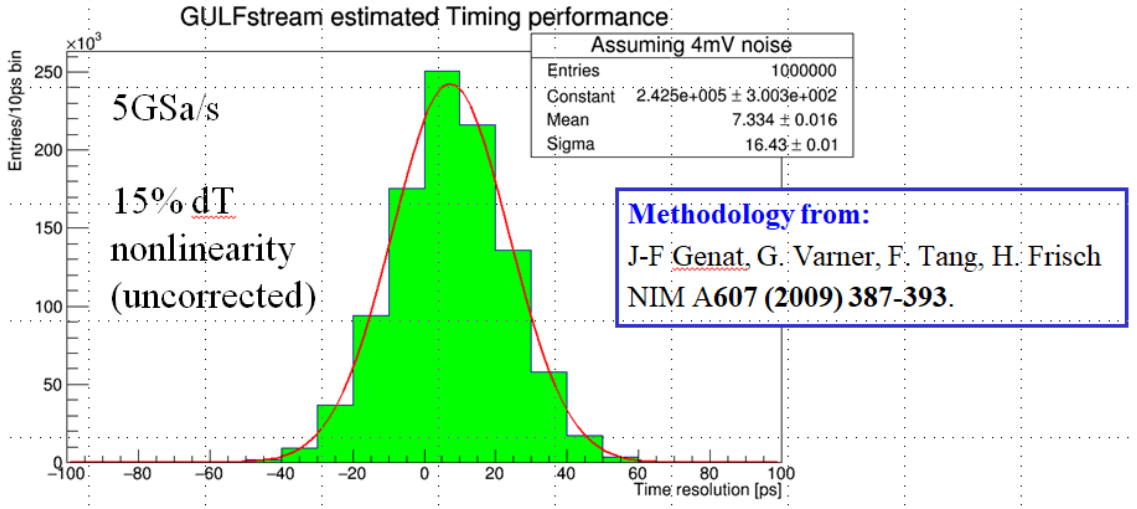
GULFstream Operating Principle

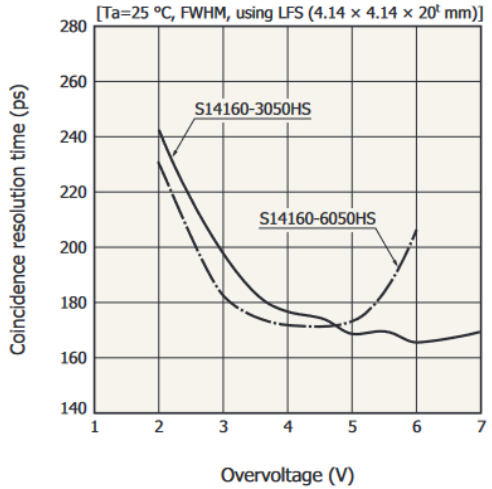
- For MCPs, charge distribution goes all the way to zero
- For SiPMs, quantized; Time to Digital Converter methodology makes sense
- SCA sampling at ~ 4GSa/s; trigger on stored samples; priority encoder time bin
- Target interpolation: 1ns risetime, 0.5V, 64 bit sampling



1.8V process







180ps FWHM ~ 76ps

If 20ps due to ASIC (FTSW limited?)

76ps ➔ 79ps

A summary of ASIC design @ UH

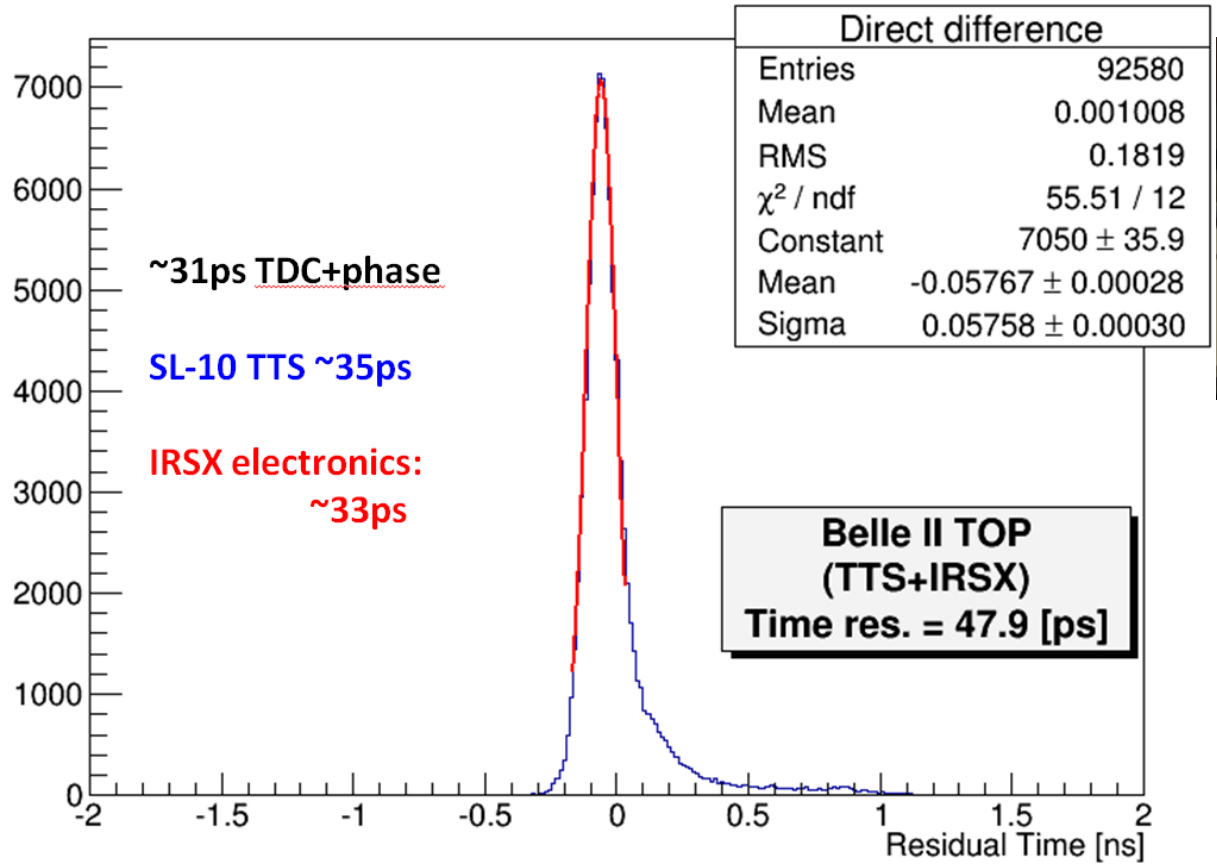
- **High impact from a tiny group; physically the most isolated in the world**
- **Didn't mention all of the endeavors (such as Q-Pix, GRAPH, …)**
  - ❖ **(nor any CAP (MAPS) pixel activities)**
- **Students and Postdocs have made a huge impact**

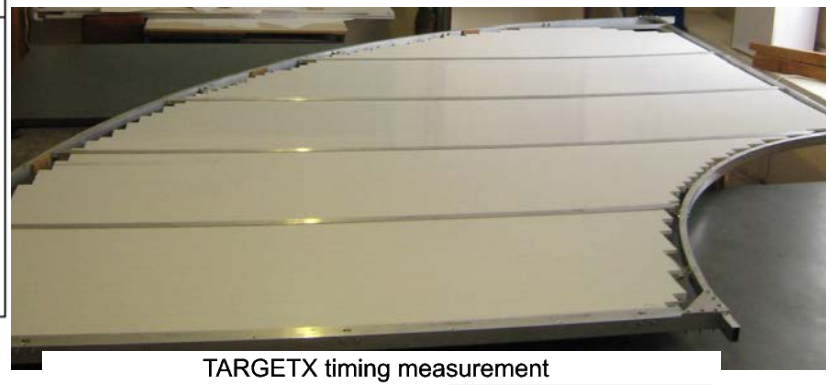
- **The latter also realized the goal of commercializing the technology (Nalu Scientific)**

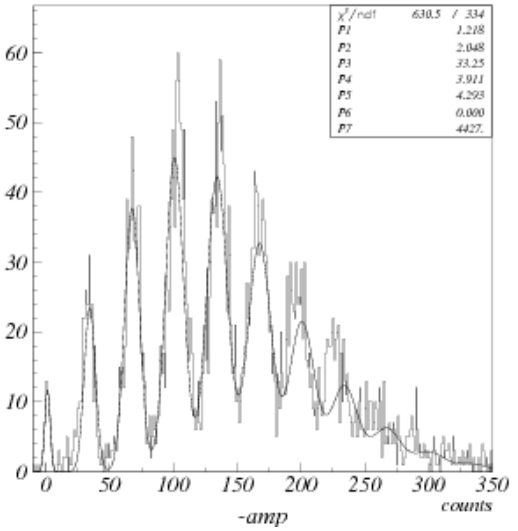
- **Opportunities: we expect to be recruiting for an Instrumentation Frontier faculty position soon**

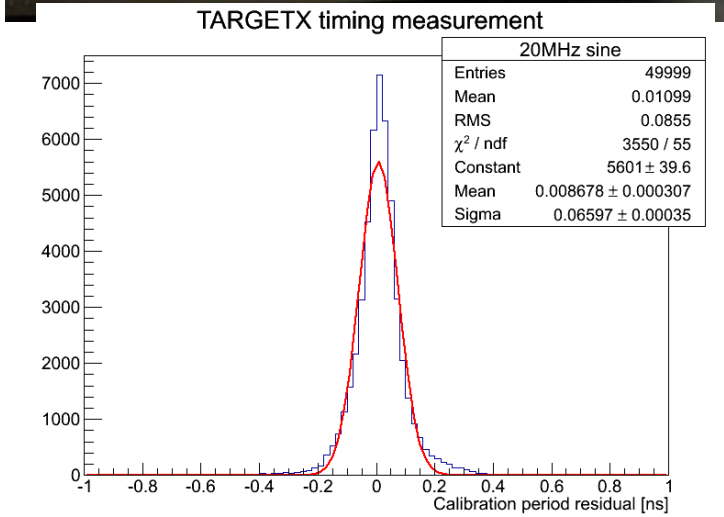
Belle II iTOP and KLM: IRSX, TARGETX

- **Leveraged decade of development for PID and Muon readout**









< 0.1 ns

- **IRSX 8-channels, 32k samples/channel**
- **8k channels of MCP-PMT readout**
- **30kHz L1 trigger rate (5.2us L1 delay)**

- **TARGETX 16-channels, 16k samples/channel**
- **23k channels of Si-PMT readout**
- **$1.40/channel**